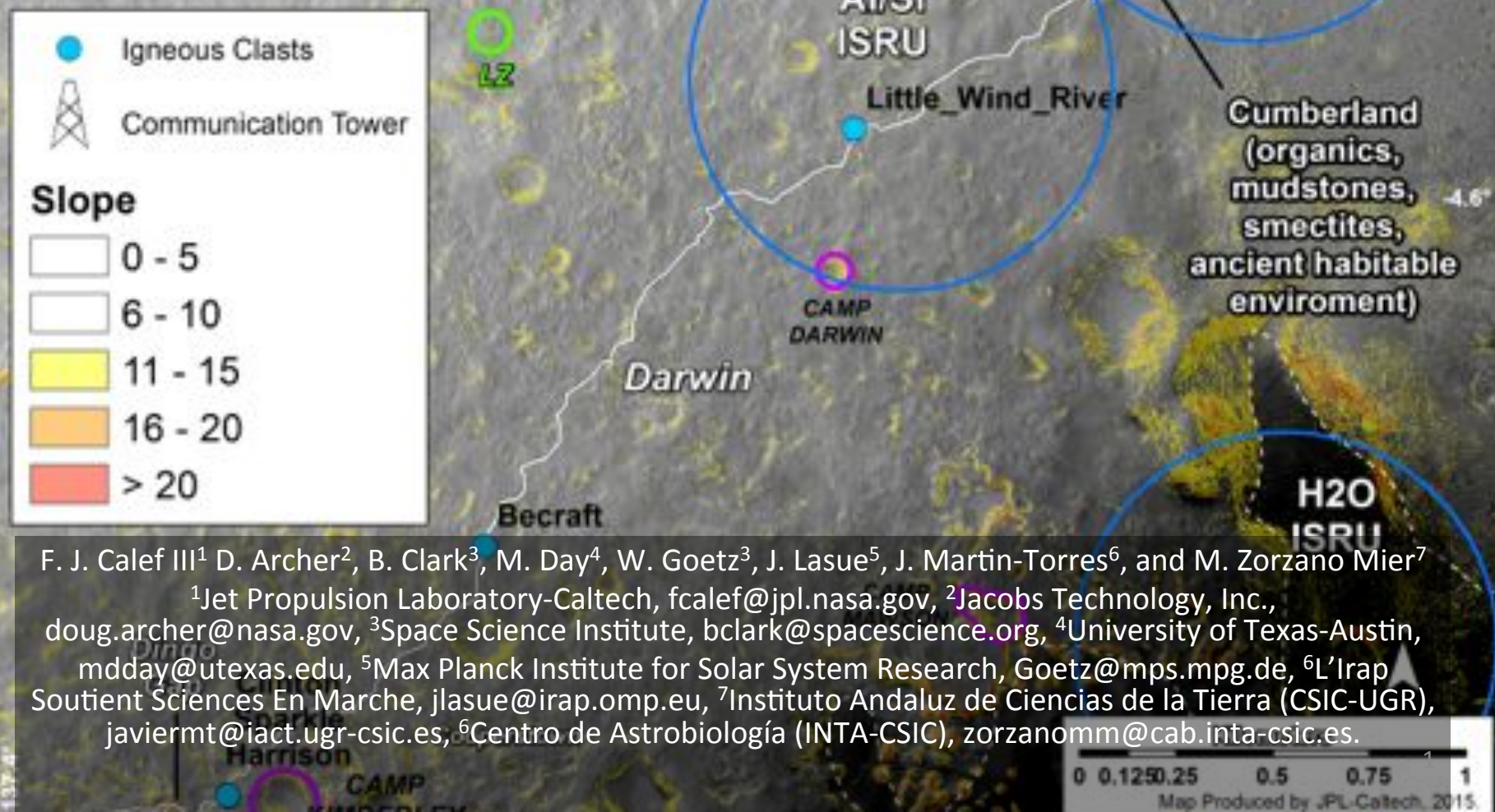


# ASSESSING GALE CRATER AS A POTENTIAL HUMAN MISSION LANDING SITE ON MARS (#1020)



F. J. Calef III<sup>1</sup>, D. Archer<sup>2</sup>, B. Clark<sup>3</sup>, M. Day<sup>4</sup>, W. Goetz<sup>3</sup>, J. Lasue<sup>5</sup>, J. Martin-Torres<sup>6</sup>, and M. Zorzano Mier<sup>7</sup>

<sup>1</sup>Jet Propulsion Laboratory-Caltech, [fcalef@jpl.nasa.gov](mailto:fcalef@jpl.nasa.gov), <sup>2</sup>Jacobs Technology, Inc., [doug.archer@nasa.gov](mailto:doug.archer@nasa.gov), <sup>3</sup>Space Science Institute, [bclark@spacescience.org](mailto:bclark@spacescience.org), <sup>4</sup>University of Texas-Austin, [mdday@utexas.edu](mailto:mdday@utexas.edu), <sup>5</sup>Max Planck Institute for Solar System Research, [Goetz@mps.mpg.de](mailto:Goetz@mps.mpg.de), <sup>6</sup>L'Irap Soutient Sciences En Marche, [jasue@irap.omp.eu](mailto:jasue@irap.omp.eu), <sup>7</sup>Instituto Andaluz de Ciencias de la Tierra (CSIC-UGR), [javiermt@iact.ugr-csic.es](mailto:javiermt@iact.ugr-csic.es), <sup>6</sup>Centro de Astrobiología (INTA-CSIC), [zorzanomm@cab.inta-csic.es](mailto:zorzanomm@cab.inta-csic.es).



# “Go Where You Know”

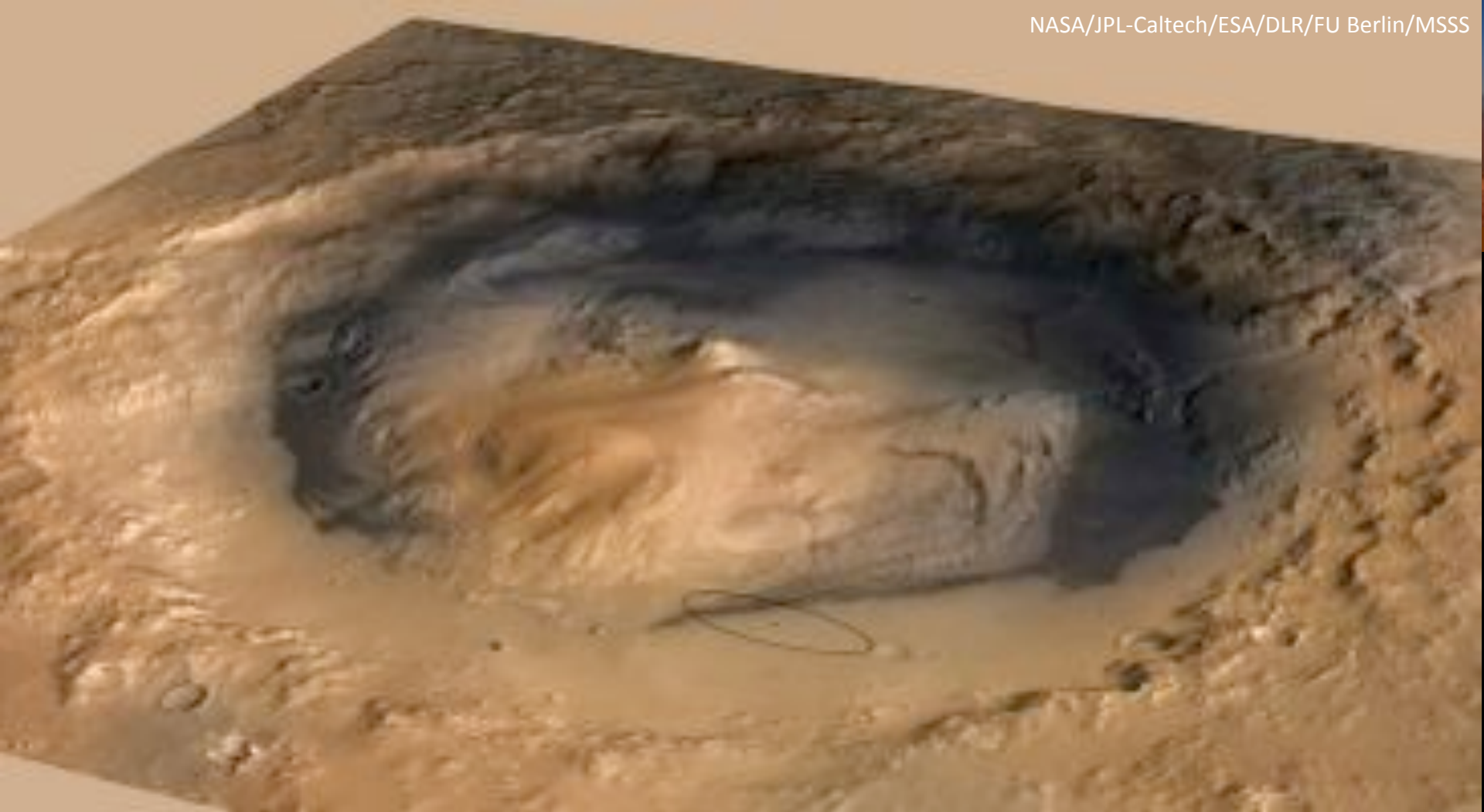
1<sup>st</sup> EZ Workshop for Human Missions to Mars

Three low-latitude sites with extensive ground truth exist: *Meridiani Planum*, *Gusev Crater*, and ***Gale Crater***; they offer steady climatic conditions, cm-scale hazard assessments, and well-characterized science regions of interest (ROIs).

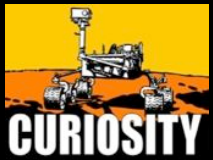


This presentation aims to show why ***Gale Crater*** offers several compelling science targets and quantified ISRU resources based on insitu observations measured from the unique set of instruments onboard the Mars Science Laboratory (MSL) rover mission.





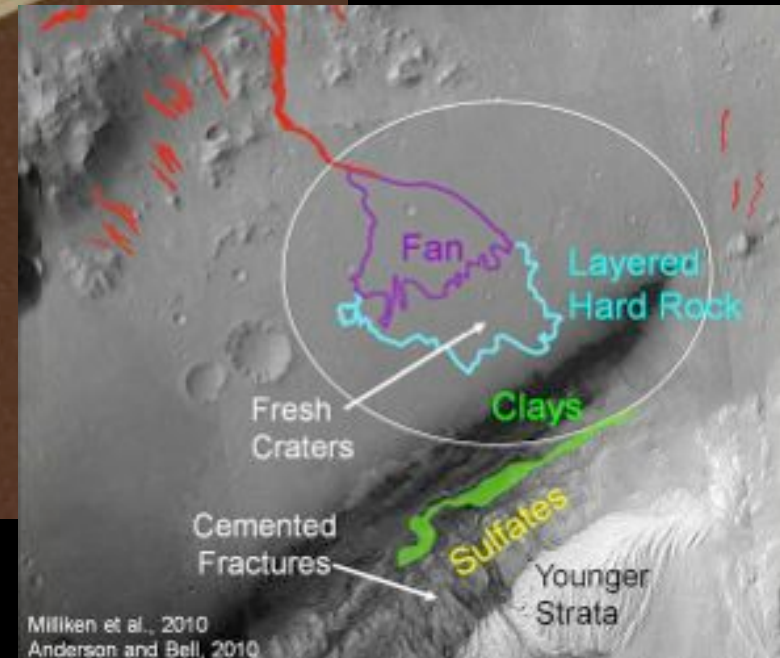
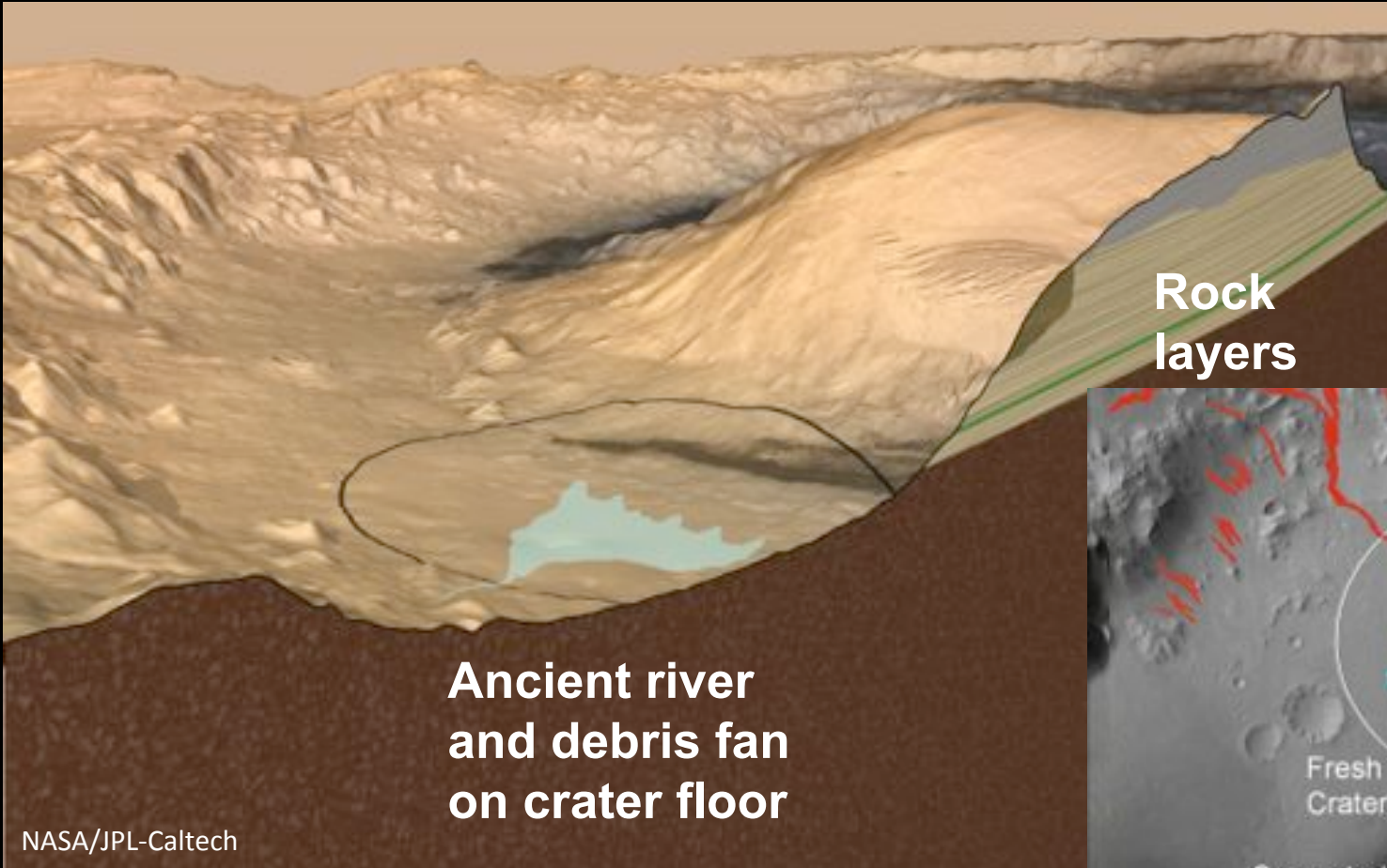
**155-km Gale Crater contains a 5-km high mound of stratified rock. Strata in the lower section of the mound are composed of clays and sulfates, while the upper mound is dry, suggesting transition from 'wet' Mars to 'dry' Mars (Late Noachian to Early Hesperian?).**





# Water-Related Geology and Minerals at Mount Sharp: a 5 km Stratigraphic Record of Mars' Past

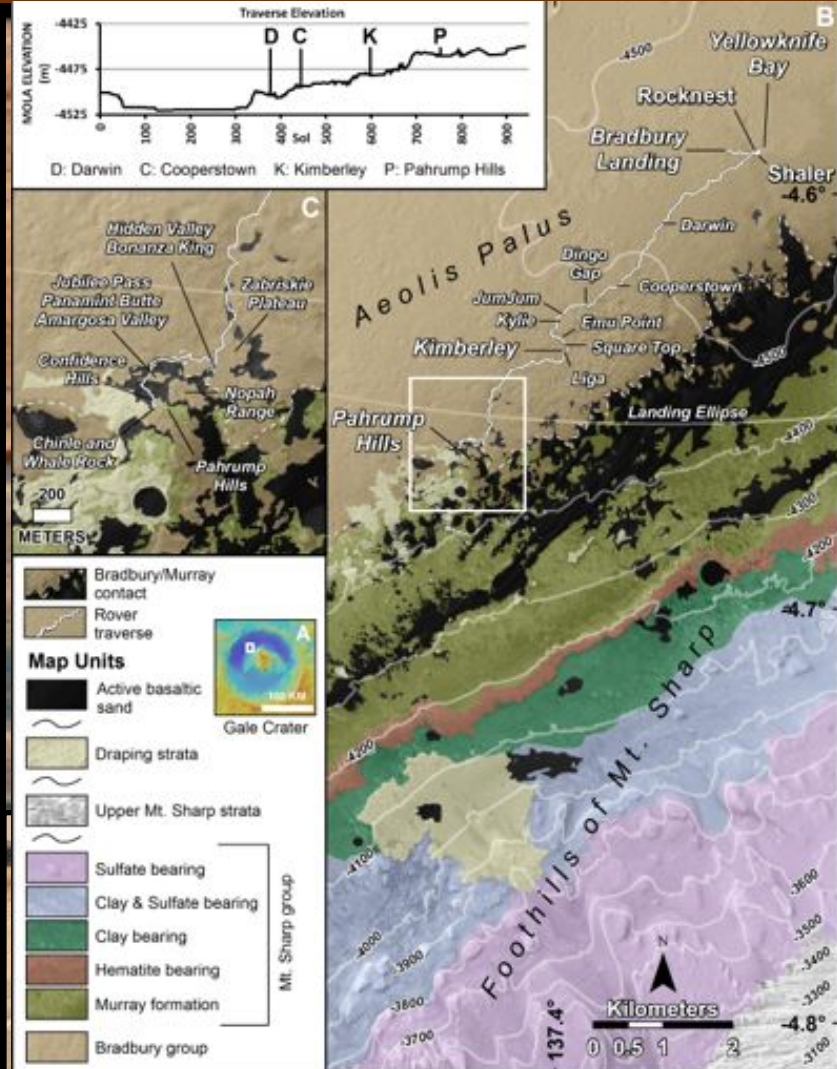
1<sup>st</sup> EZ Workshop for Human Missions to Mars





# Aeolis Mons (Mt. Sharp) Geology: Transition from 'Wet' to 'Dry' Mars

1<sup>st</sup> EZ Workshop for Human Missions to Mars



'Dust-stone' Unit (~10 km)

Sulfate Unit (8 km)

Clay Unit (6 km)

Hematite Ridge (5 km)

Paintbrush Unit (2 km)

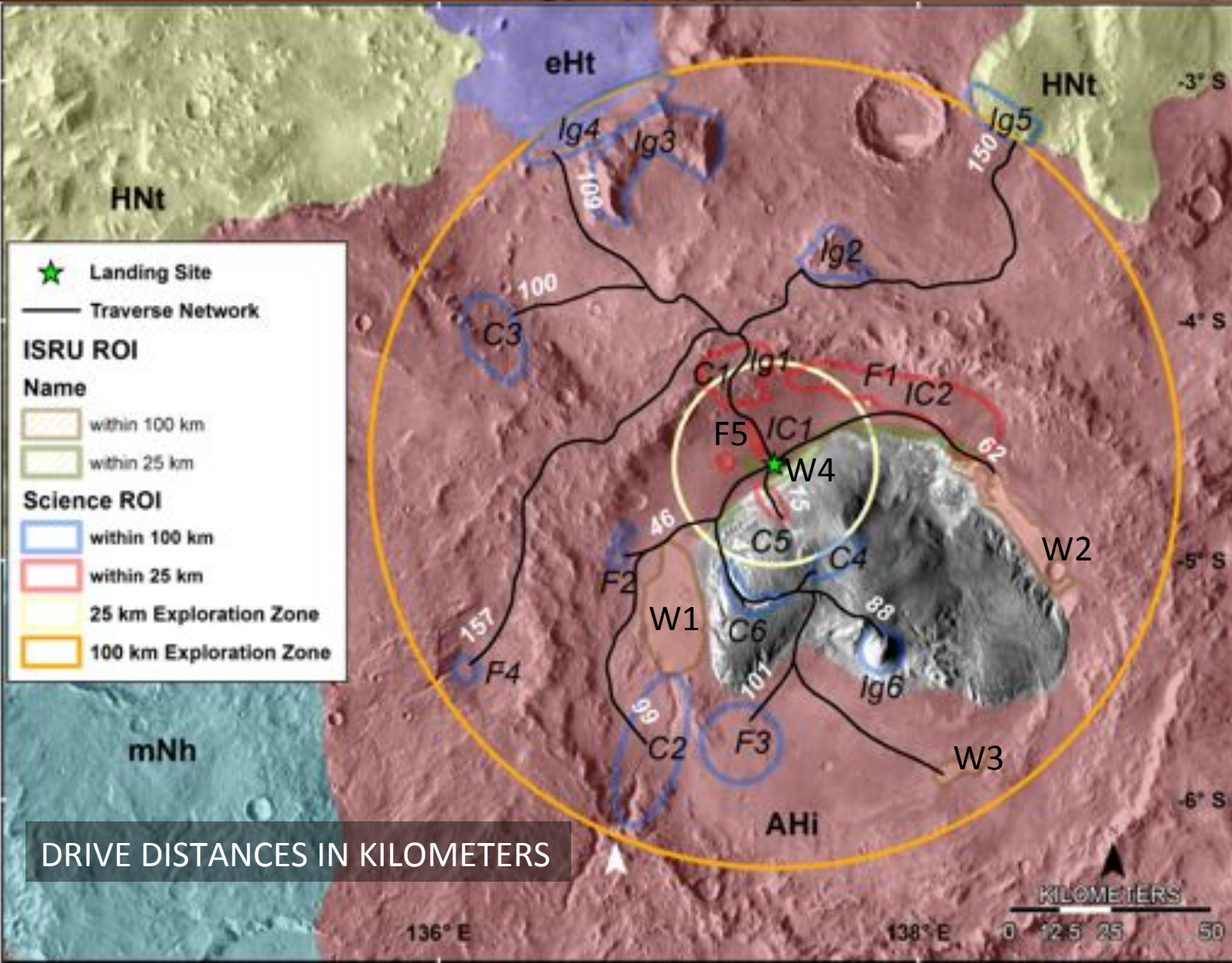
Bagnold Dunes

Left figure from Grotzinger et al., Science, 2015

NASA/JPL-Caltech/MSSS



# Gale Crater Exploration Zone



Human Missions to Mars

## LZ Coordinates

Longitude (E)

137.42009295°

Latitude (S)

4.59310427°

Elevation (MOLA)

-4497.77

Easting

8145534.27 m

Northing

-272254.70 m

## Science ROI

C = Channel

F = Fan

Ig = Igneous

IC = Inverted  
Channels

## ISRU ROI

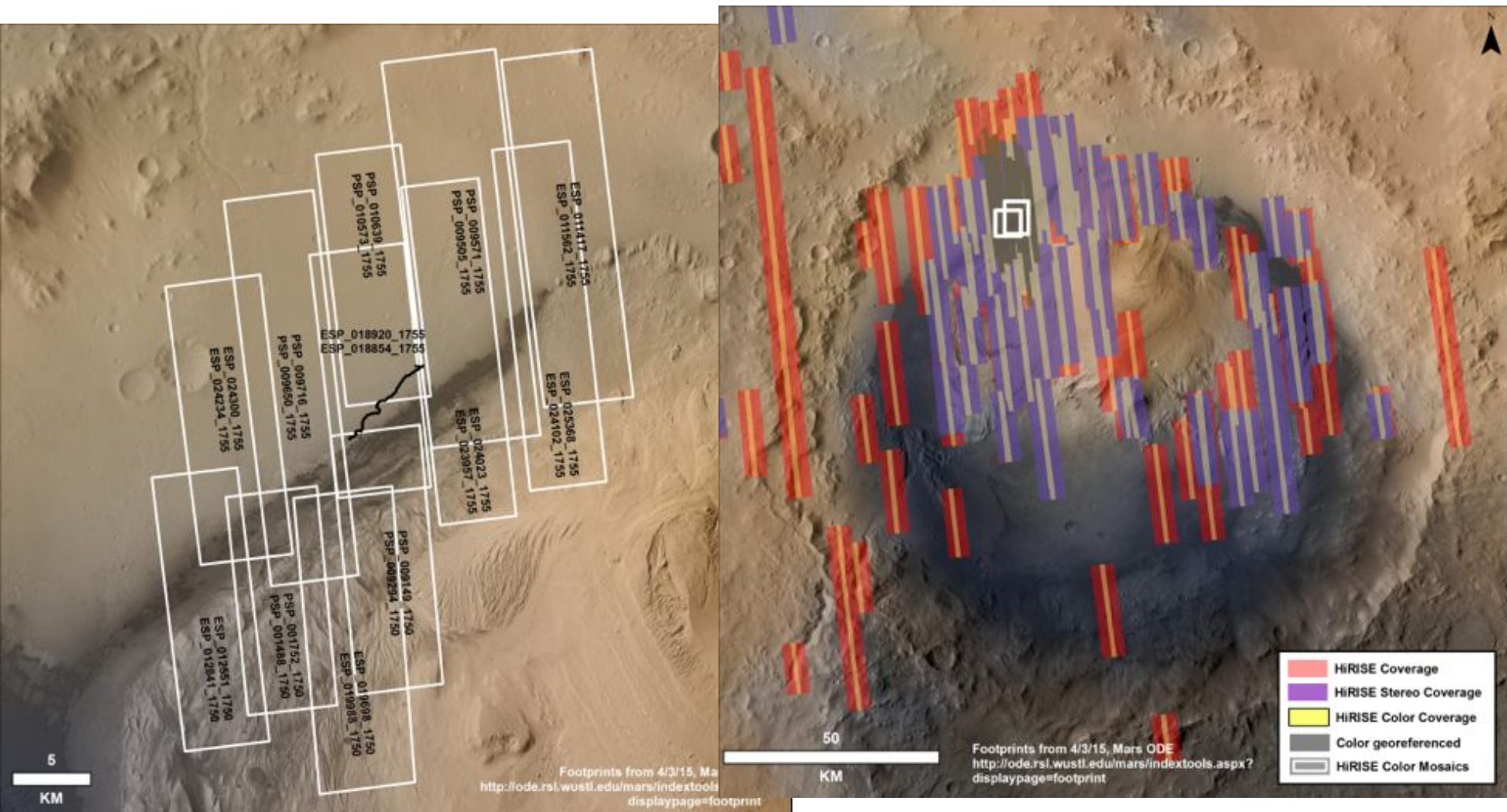
W = Water ROI



# Extensive HiRISE Coverage for Analyzing Science and ISRU ROIs



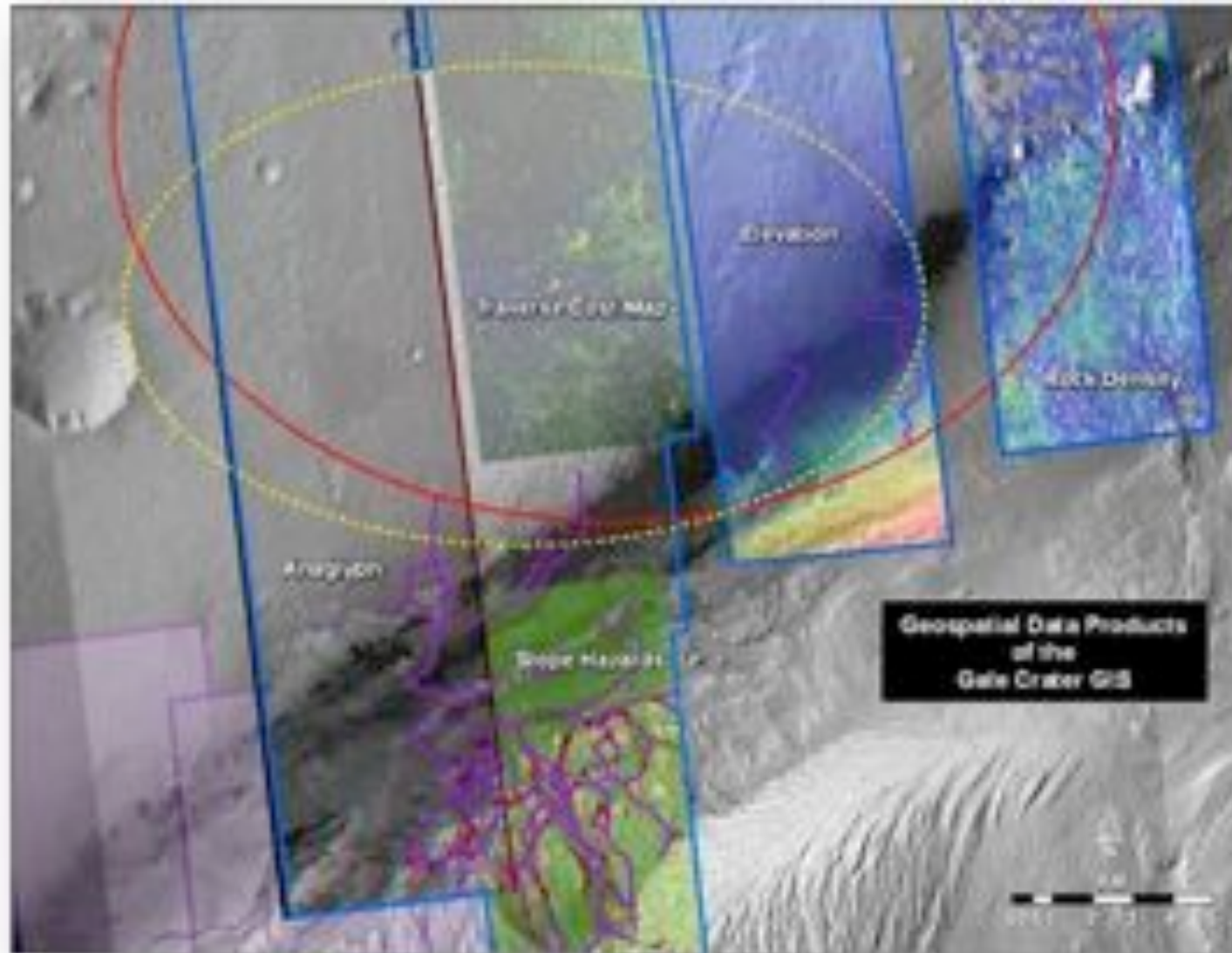
1<sup>st</sup> EZ Workshop for Human Missions to Mars





# *Multiple Co-Registered Datasets in Landing Zone: Slope, Rock Abundance, Hazards, and more*

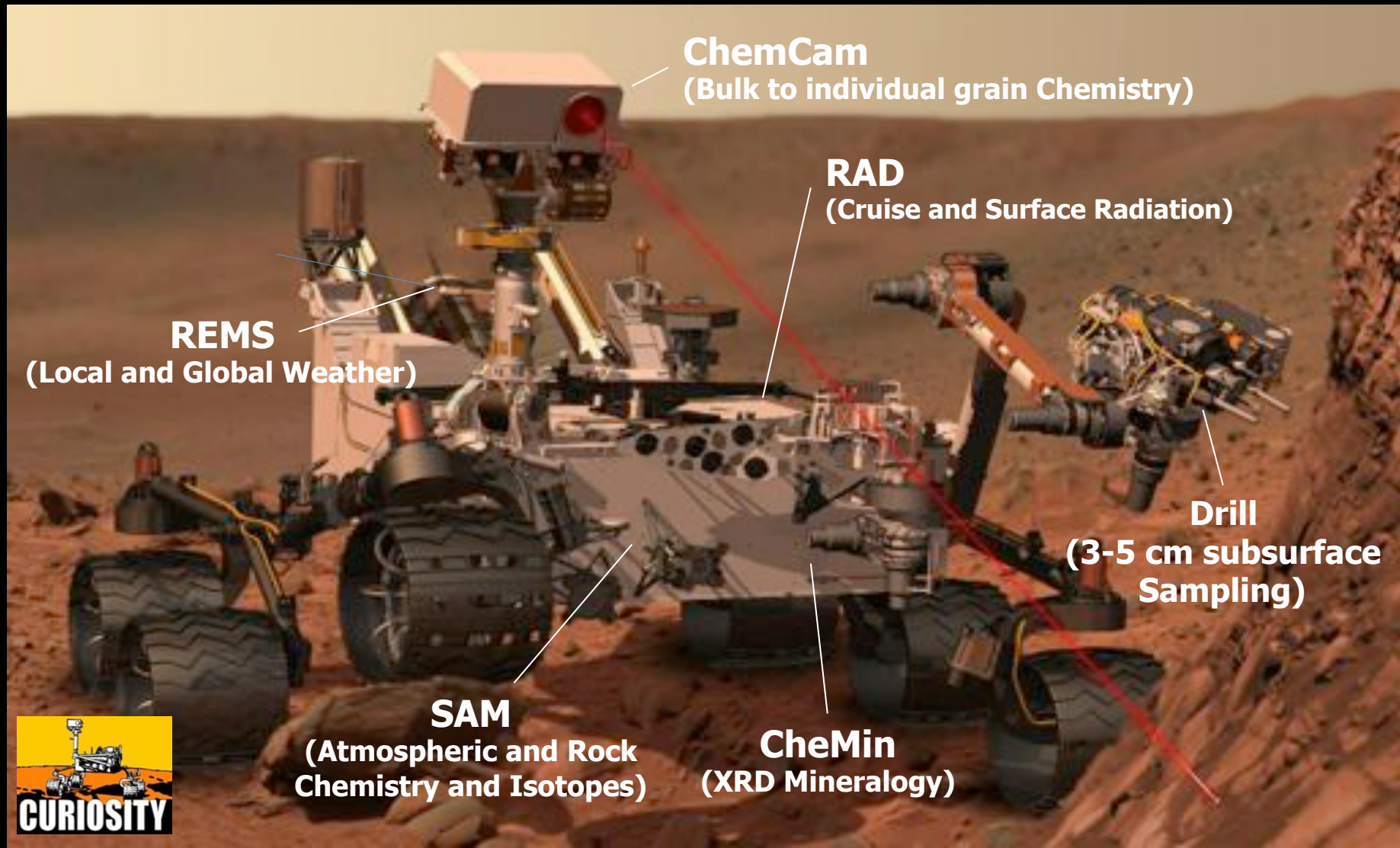
1<sup>st</sup> EZ Workshop for Human Missions to Mars





# Curiosity's Science Payload Allows Unique Science & ISRU ROI Characterization

1<sup>st</sup> EZ Workshop for Human Missions to Mars



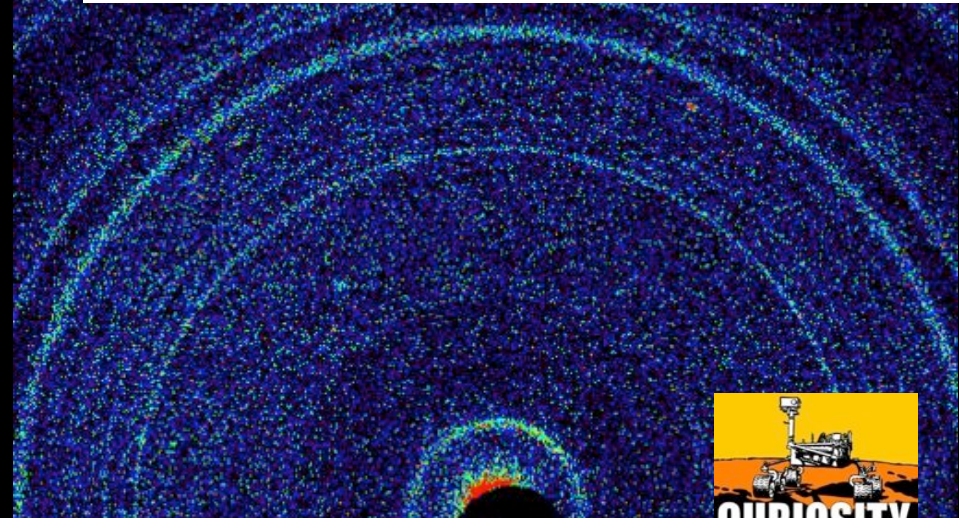
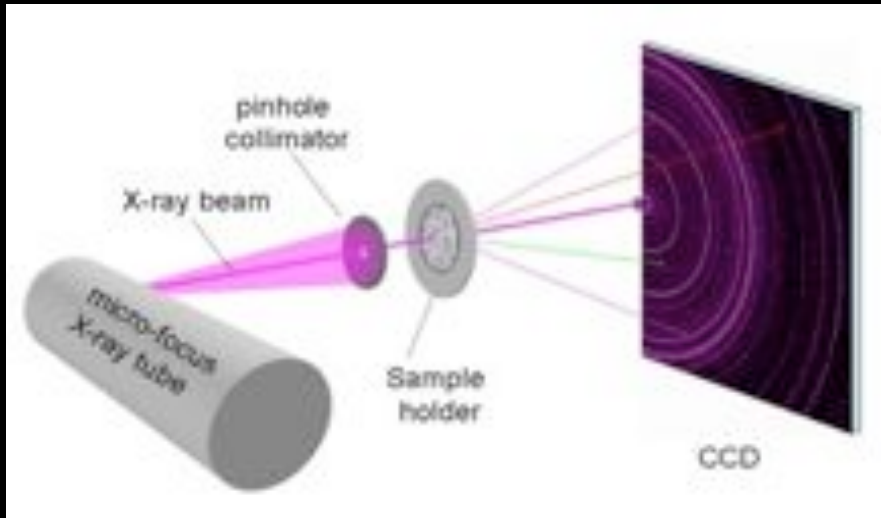
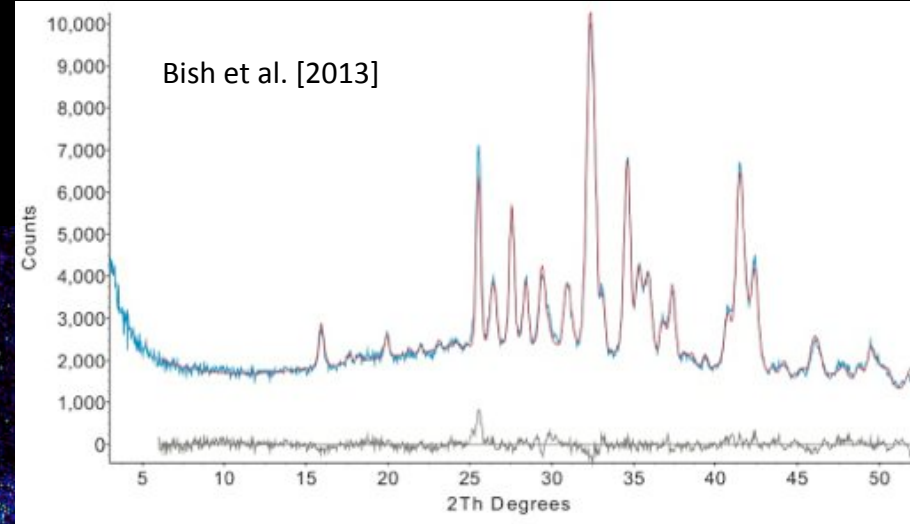


**Rocknest sand has a typical Mars basalt composition, but also 1.5-3% bound water. Both a water and food growth medium.**

1<sup>st</sup> EZ Workshop for Human Missions to Mars

**Rocknest sand is composed of unaltered basaltic minerals, typical of rocks and soils on Mars**

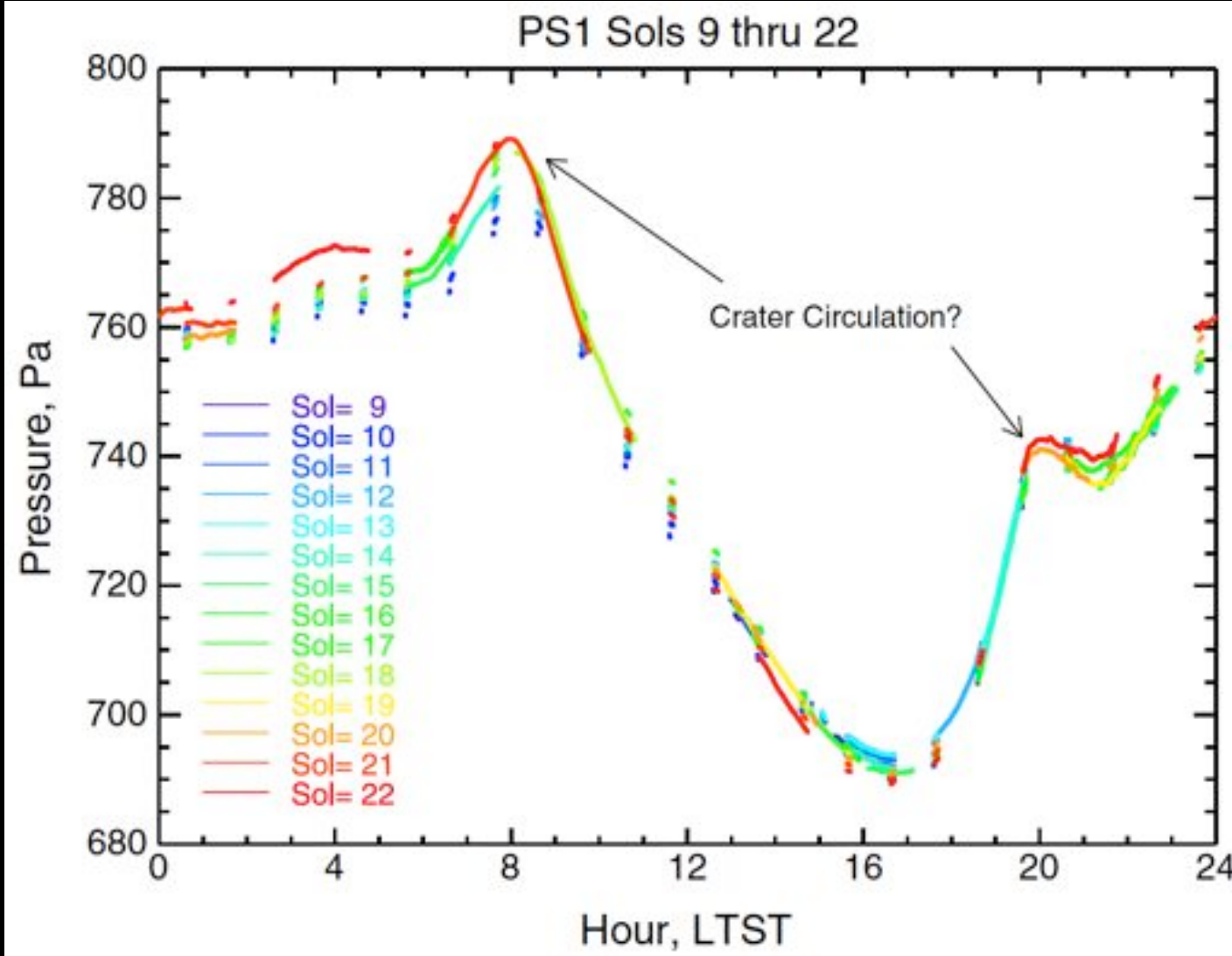
**Samples also contained 30-45% non-crystalline material, containing volatiles, sulfur, calcium, and perhaps nanophase ferric oxide.**





# REMS insitu pressure measurements reveal local, mesoscale, and planetary phenomena

1<sup>st</sup> EZ Workshop for Human Missions to Mars



**REMS takes hourly measurements with occasional 1-Hz extended sessions**

**Daytime convective vortices are present, but no dust devils have been observed**

**Diurnal thermal tides (left) are amplified and modified by the crater topography**

**The CO<sub>2</sub> pressure cycle at Gale also has components due to elevation and planetary circulations**





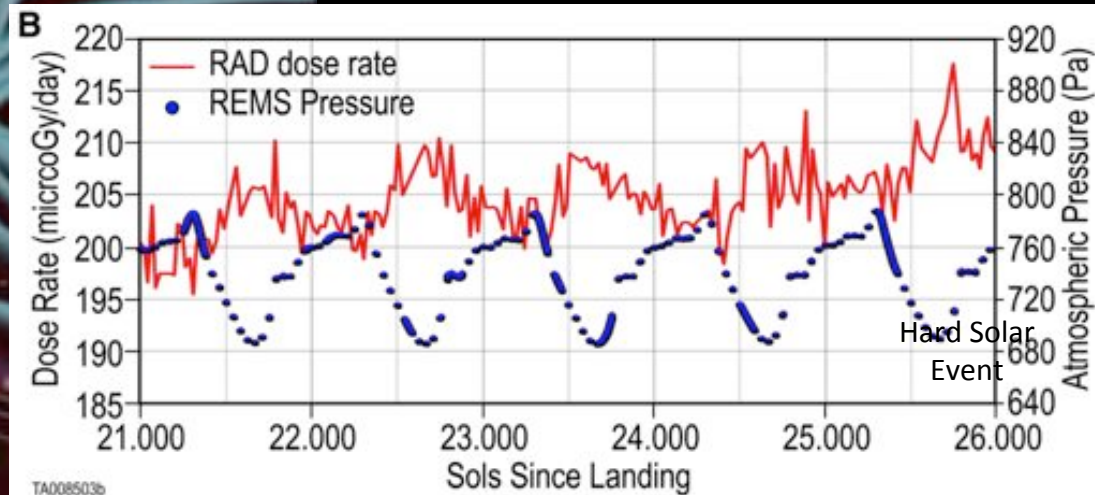
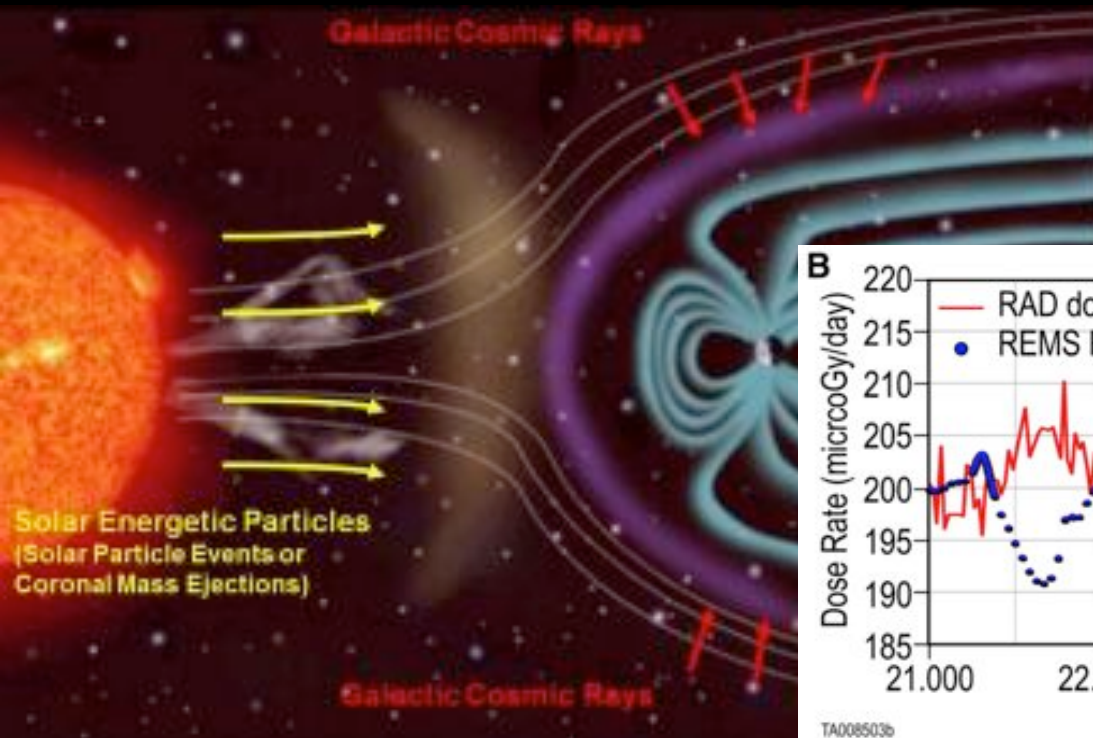
The RAD instrument measured the radiation flux from both galactic cosmic rays and solar energetic particles:  
*measured radiation hazard at the surface*

1<sup>st</sup> EZ Workshop for Human Missions to Mars

The surface dose rate is about half that measured in cruise

A crewed mission would receive ~1 Sievert of exposure in a trip to Mars with 500 sols on the surface

[Hassler et al., 2014]



Curiosity's Radiation Assessment Detector measures high-energy radiation



# Atmospheric Gas Abundances Measured by SAM: *Methane detection on surface*

1<sup>st</sup> EZ Workshop for Human Missions to Mars

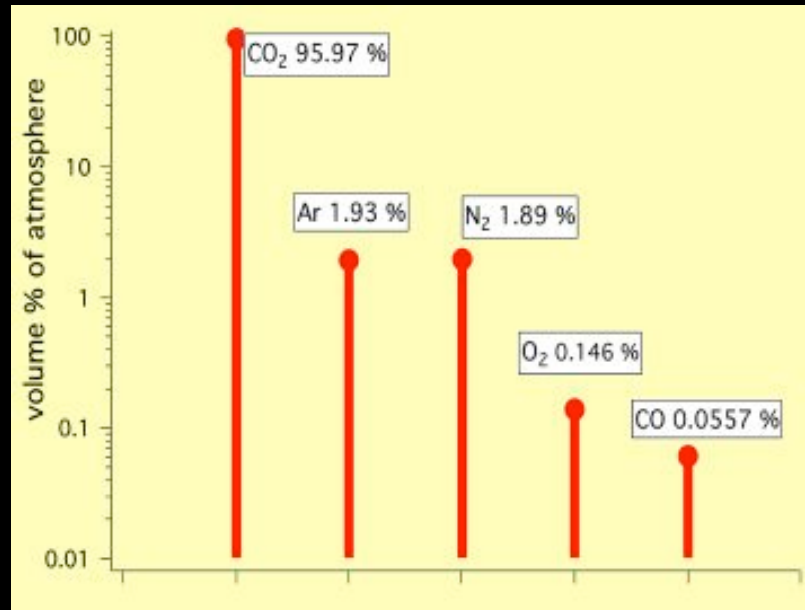
SAM also found that Mars' atmosphere is enriched in the heavy versions of isotopes, indicating massive atmospheric loss to space

$\delta^{13}\text{C} = 46 \pm 4$  per mil

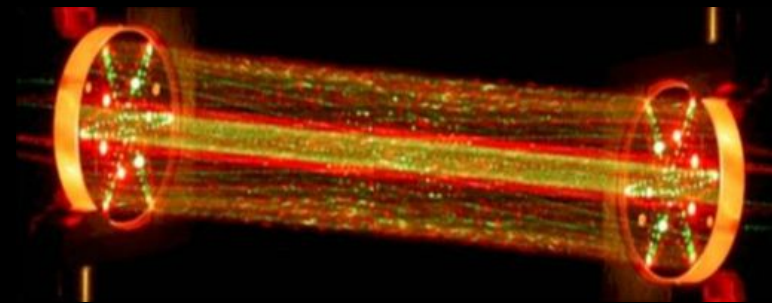
$\delta\text{D} = 4950 \pm 1080$  per mil

$^{40}\text{Ar}/^{36}\text{Ar} = 1900 \pm 300$

Methane HAS been definitively detected  
Upper limit = ~7 ppb

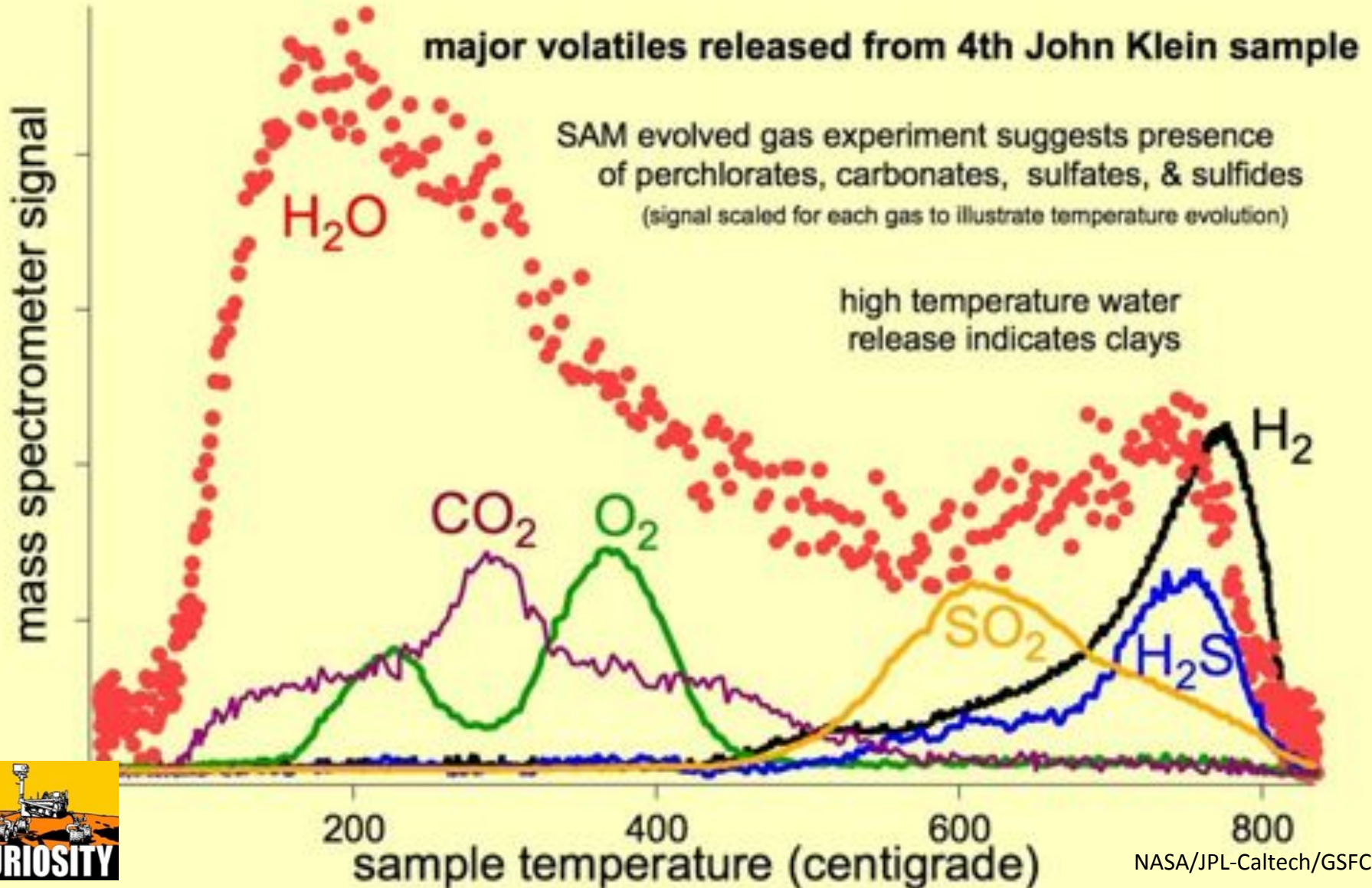


SAM found that argon, rather than nitrogen is the second most abundant gas





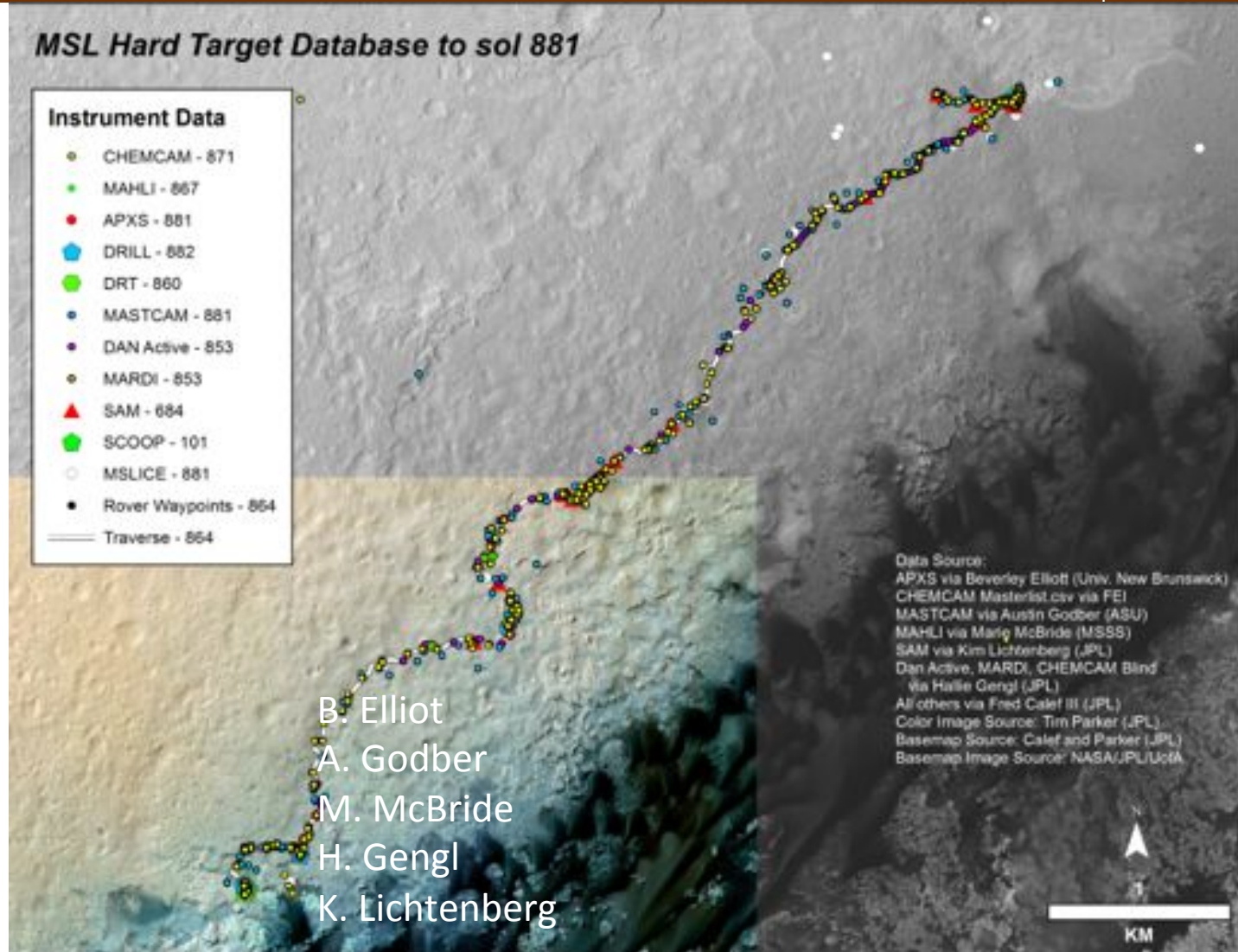
# Major gases released from John Klein sample and analyzed by SAM: determined Yellowknife Bay was a Habitable Environment when Gale Crater had a Lake





# Tracking Science Observations: thousands of potential sample locations known to cm scale

1<sup>st</sup> EZ Workshop for Human Missions to Mars

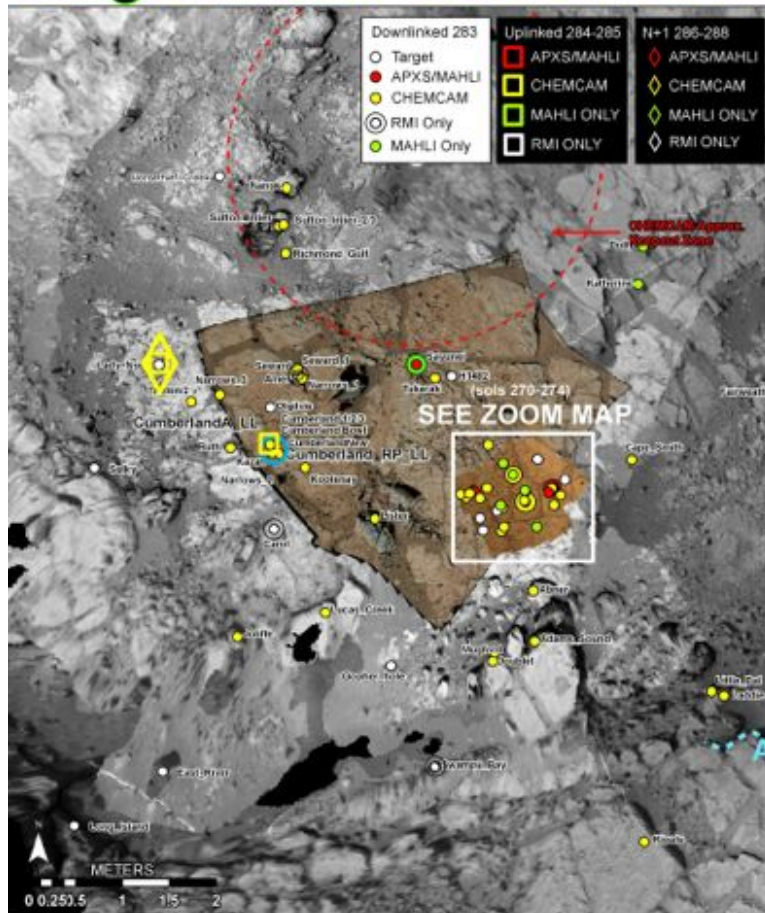




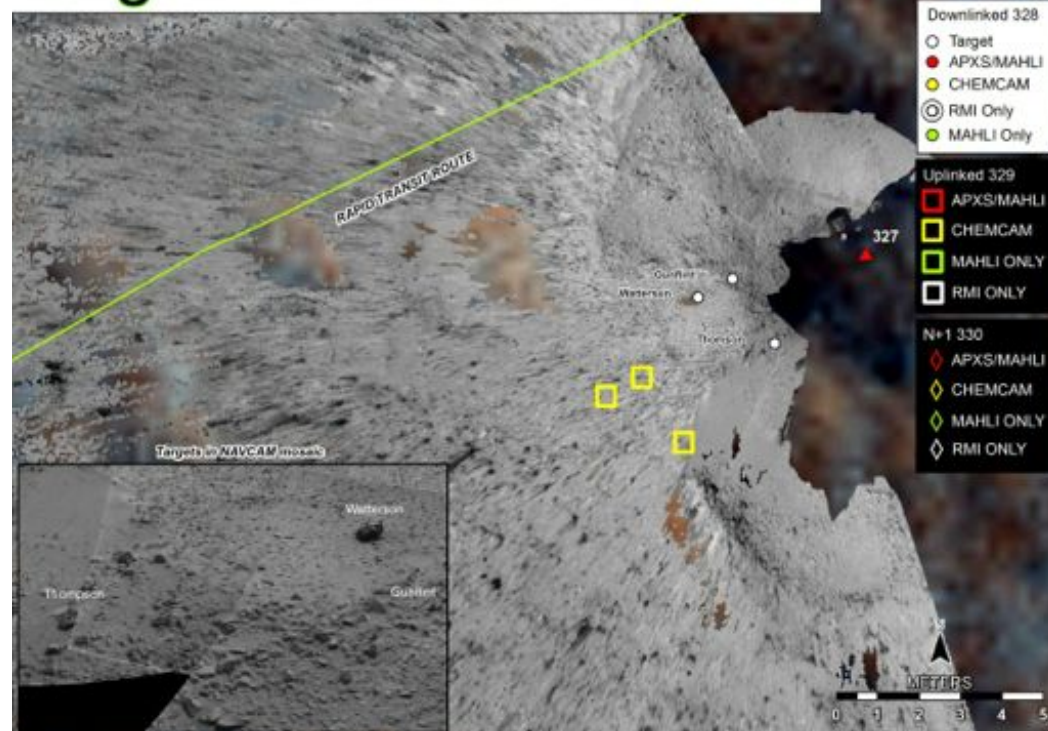
# Example Target Map from MSL Operations: known locations for future science opportunities maximize science return during field excursions

1<sup>st</sup> EZ Workshop for Human Missions to Mars

## Target Data: Sol 283-288



## Target Data: Sol 328-330



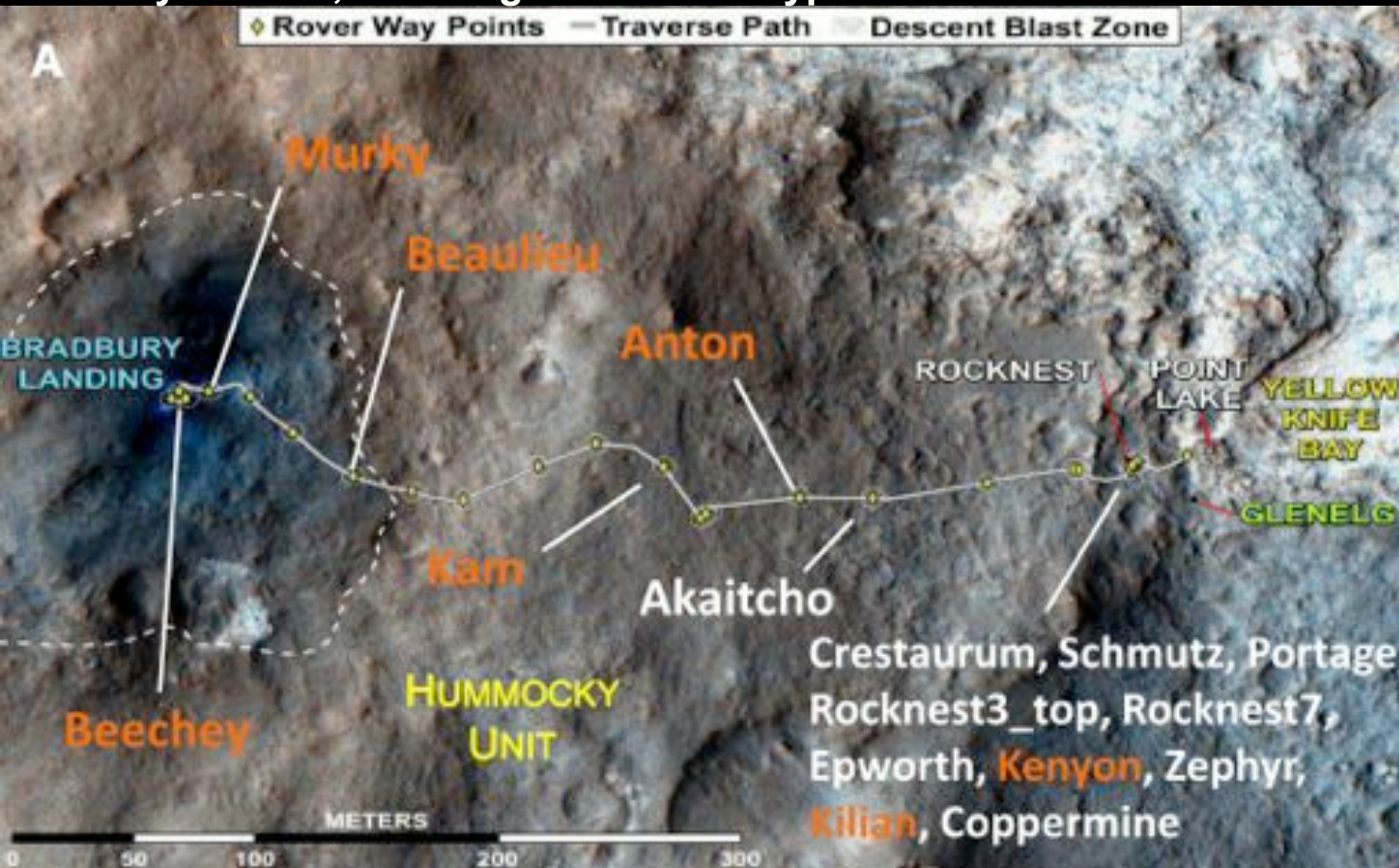


# Felsic vs. Mafic ISRU Source Areas

1<sup>st</sup> EZ Workshop for Human Missions to Mars

ChemCam identified two principal soil types along the traverse to Yellowknife Bay: a fine-grained, mafic type similar to other soils, and a locally derived, coarse-grained felsic type.

Mafic soil component has hydration signature, corresponding to the X-ray amorphous component sampled by CheMin and SAM.



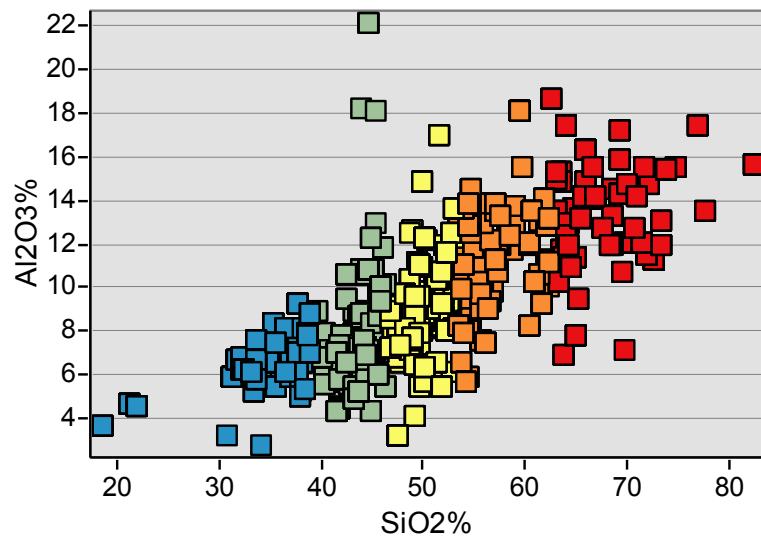
Meslin et al. [2013]

**Mafic**

**Felsic**



SiO<sub>2</sub> vs Al<sub>2</sub>O<sub>3</sub>



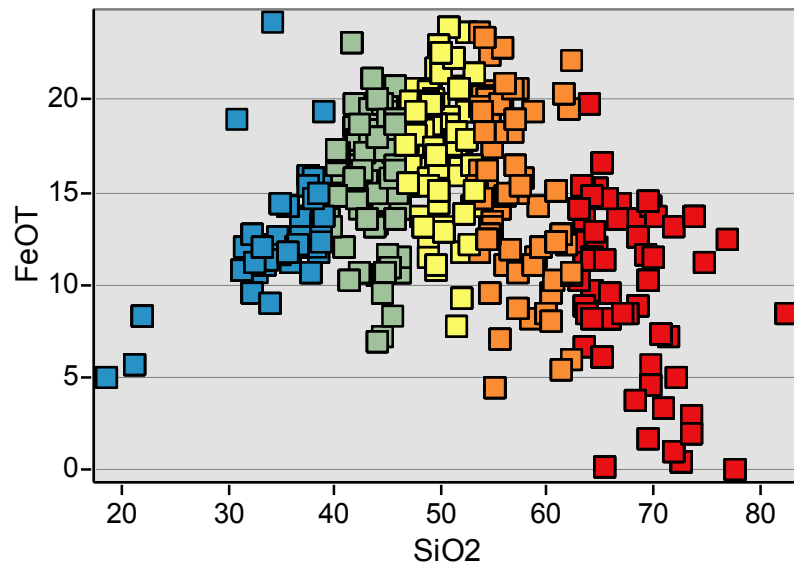
CHEMCAM Oxides

>10 % Al<sub>2</sub>

Aluminum  
Detection and  
Abundance along  
MSL Traverse



SiO<sub>2</sub> vs FeOT



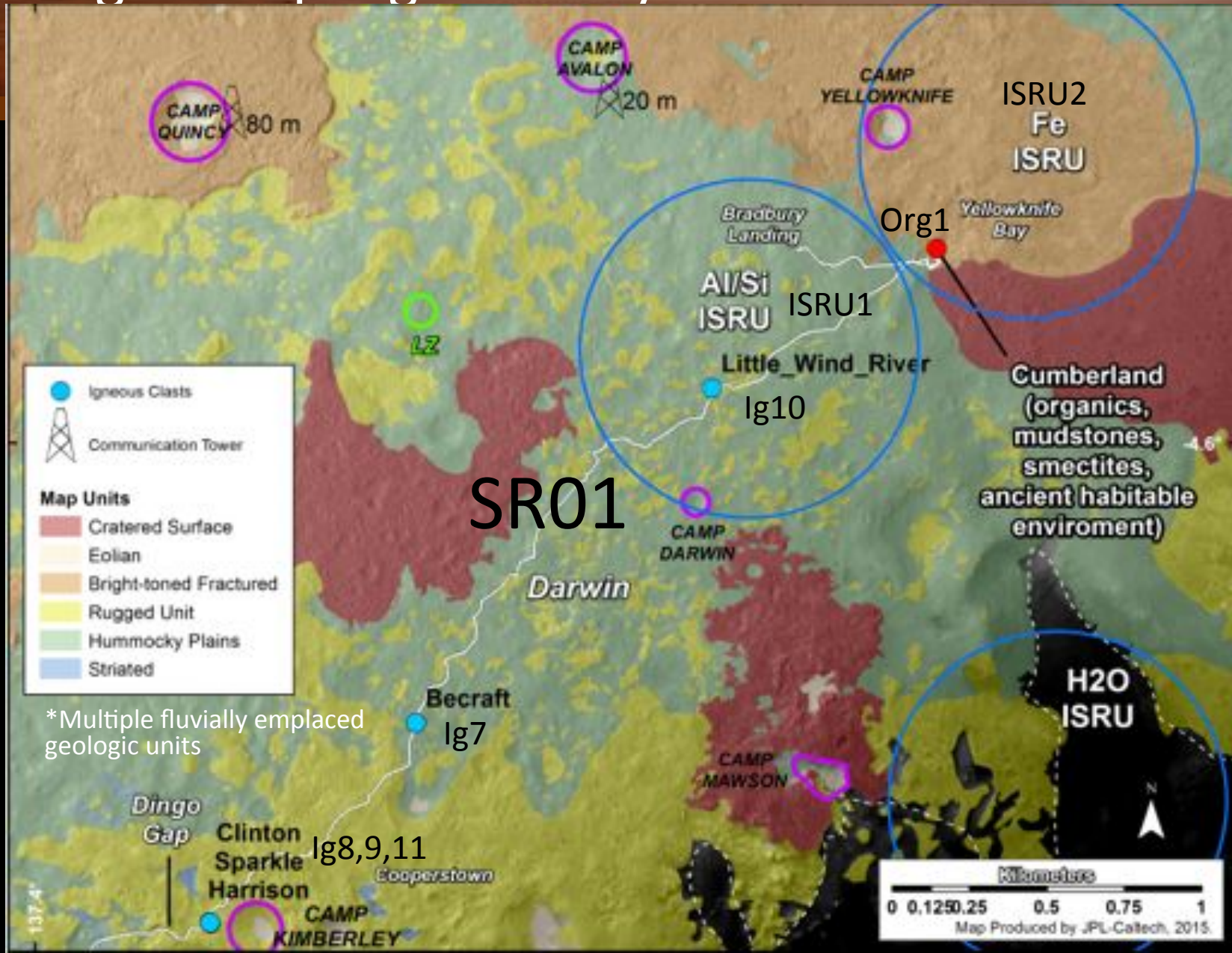
CHEMCAM\_oxides

>15 % Fe

Iron Detection and  
Abundance along  
MSL Traverse



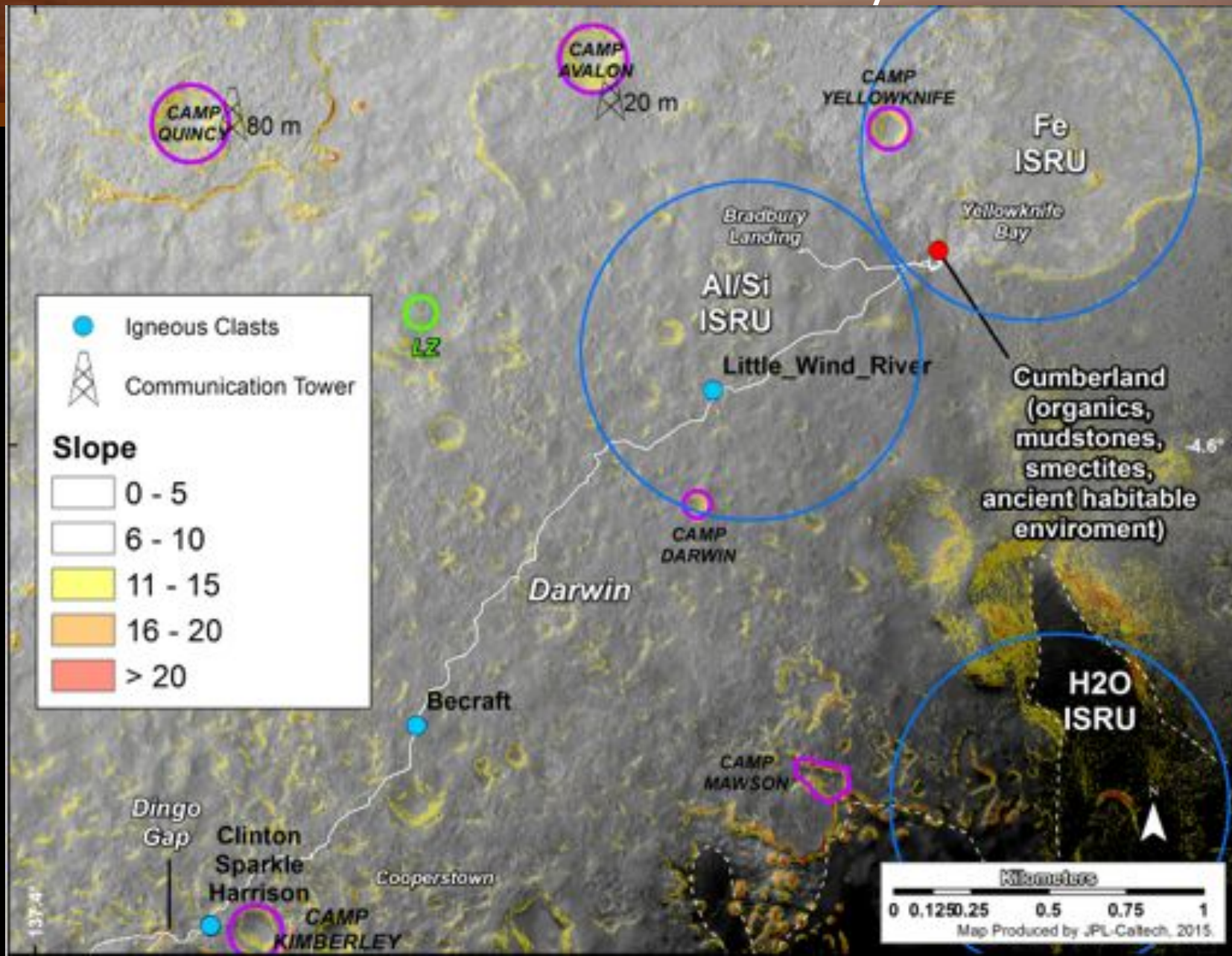
# 1<sup>st</sup> Mission Exploration Zone: Geologic Sampling Diversity within 1-2 Kilometers



ons to Mars



# 1<sup>st</sup> Mission Exploration Zone: Potential Habitation Sites and nearby ISRU Zones



ns to Mars



# ISRU Resource Numbers: Water

1<sup>st</sup> EZ Workshop for Human Missions to Mars

- Water
  - Based on Leshin et al. (2013), Rocknest sands contained up to 3 wt% H<sub>2</sub>O via SAM, Bagnold Dune field contains 10<sup>4</sup>-10<sup>6</sup> MT **adsorbed** water.
  - Easy to process. Potentially ‘reuseable’.



# ISRU Resource Numbers: Metals

1<sup>st</sup> EZ Workshop for Human Missions to Mars

- Aluminum and Silicon
  - ISRU01 rocks on average have ~12 wt% Al and ~55-60 wt% Si. ISRU02 rocks, ~18 wt% Fe.
  - Assuming rock density ~2.5 gm/cc:
    - 300 kg/m<sup>3</sup> Al (ISRU01)
    - 1375-1500 kg/m<sup>3</sup> Si (ISRU01)
    - 450 kg/m<sup>3</sup> Fe (ISRU02)



# What the surface actually looks like at the (human) landing site...



NASA/JPL-Caltech/MSSS  
Rover from 'The Martian'





# Building material and ISRU Al/Si Resource

## Rounded pebbles and sand in the conglomerate

### “Link”

1<sup>st</sup> EZ Workshop for Human Missions to Mars



NASA/JPL-Caltech/MSSS



# Science Target: Large feldspar-rich crystals (phenocrysts) in an igneous clast (Harrison): good for geologic dating

SRIg09 Harrison (lat, lon, elev)

-4.62617745, 137.40955945, -4489.114 m

1<sup>st</sup> EZ Workshop for Human Missions to Mars





# Science ROI: Yellowknife Bay for Organics and Lake sediments (Sheepbed Mudstones)



Cumberland Drill Site with martian organic detection:  
lon., lat, elev: 137.44909593, -4.58951574, -4520.029 m



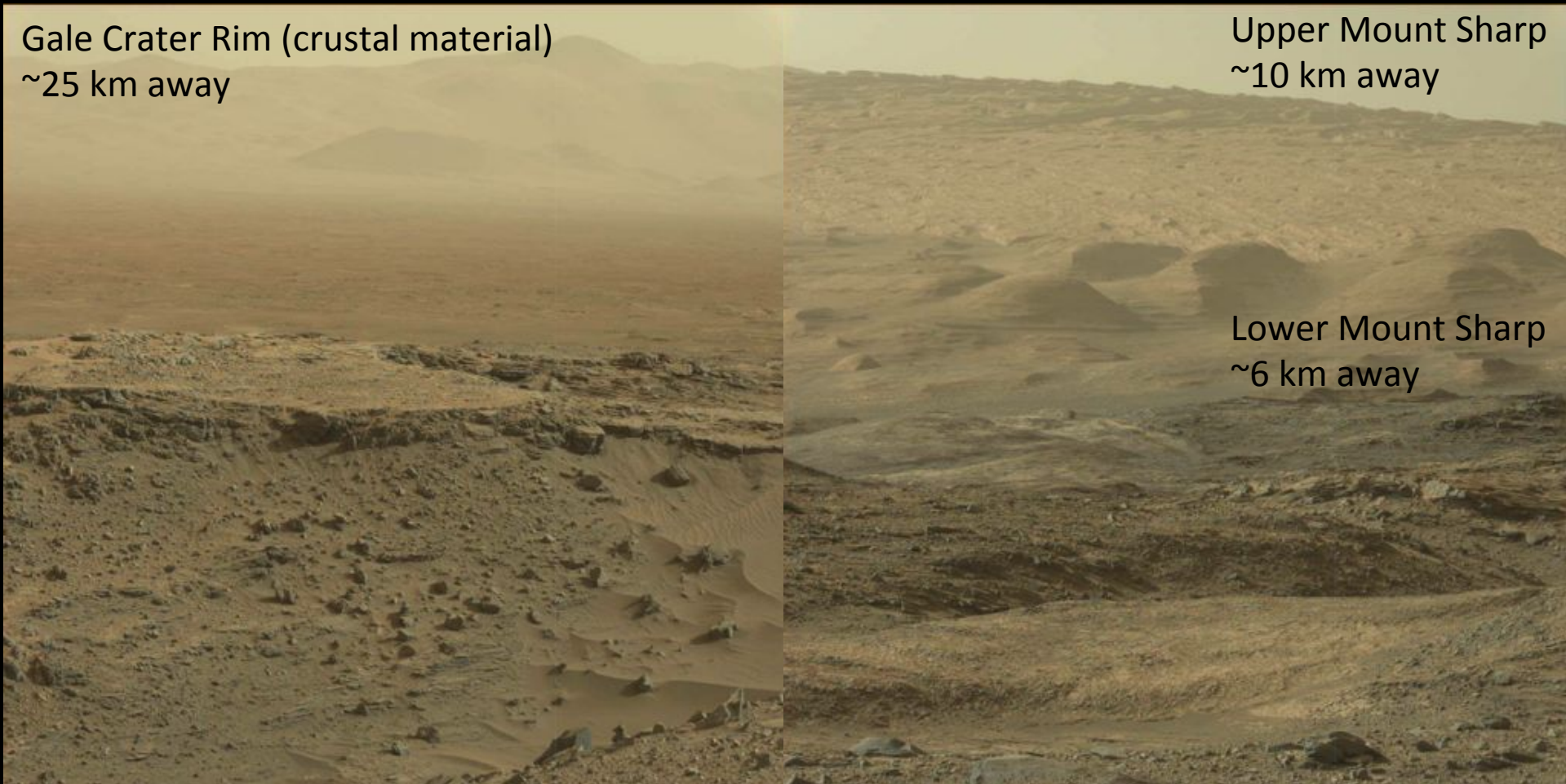


# **A Known Ancient Habitable Environment at Yellowknife Bay**

- **The regional geology and fine-grained rock suggest that the John Klein site was at the end of an ancient river system or within an intermittently wet lake bed**
- **The mineralogy indicates sustained interaction with liquid water that was not too acidic or alkaline, and low salinity. Furthermore, conditions were not strongly oxidizing.**
- **Key chemical ingredients for life are present, such as carbon, hydrogen, nitrogen, oxygen, phosphorus, and sulfur**
- **The presence of minerals in various states of oxidation would provide a source of energy for primitive organisms**



# How Close are Distal Science ROIs from Landing Site: Almost all within Walkback Distance





# Reference Mission Goals

1<sup>st</sup> EZ Workshop for Human Missions to Mars

- In regards to Mars design reference architecture.
  - Gale crater easily meets all the science and engineering goals given its selection for the MSL rover mission.
  - “Short-stay” with 30 days on the surface (500-650 days total) and “long-stay” with 500+ days on Mars (~900 days total) are supported as we have the localization of science targets down to the centimeter in a concentrated area (<10 km radius circle). \*Supports JPL ‘minimal architecture (Price, Baker, and Naderi, *New Space*, 2015).
  - Gale crater offers crustal material within ~20-25 km at the rim, Aeolis Mons (Mt. Sharp) clay to sulfate transition within ~15 km, primary alluvial fan material <5 km and fluvially derived conglomerates and clays, some lacustrine, upon landing in many areas of the current MSL ellipse. \*Well within ‘walk-back’ distance for astronauts.
  - Many known scientific outcrops at Gale crater are <4 km radius from field station/habitat, if the human habitat were placed within the current MSL landing ellipse area.
  - Outcrops would be close, safe, and pre-characterized sufficiently by Curiosity to allow short or long-stay missions.
- **Conclusion:** Gale crater provides excellent EDL, ISRU, and science opportunities for a human-rated mission.



# RUBRICS



# Science ROI(s) Rubric

1<sup>st</sup> EZ Workshop for Human Missions to Mars

Site Factors				SRIg01-06	SRIg07-11	SRIc01-02	SROrg1	SRF01-04	SRC01-06	SR01	EZ SUM
Science Site Criteria	Astrobio	Threshold	AND/OR	Potential for past habitability			○	●	●		5,1
				Potential for present habitability/refugia						○	0,1
		Qualifying		Potential for organic matter, w/ surface exposure		○	●	○	○		1,11
	Atmospheric Science	Threshold		Noachian/Hesperian rocks w/ trapped atmospheric gases			●	●		●	6,0
										●	1,0
		Qualifying		Meteorological diversity in space and time						●	1,0
				High likelihood of surface-atmosphere exchange						●	1,0
				Amazonian subsurface or high-latitude ice or sediment							
				High likelihood of active trace gas sources						○	0,1
	Geoscience	Threshold		Range of martian geologic time; datable surfaces	●	●		●	●		21, 0
				Evidence of aqueous processes		●	●	●	●		13, 0
				Potential for interpreting relative ages	●	●	●	●	●		20, 0
		Qualifying		Igneous Rocks tied to 1+ provinces or different times	●	●					11, 0
				Near-surface ice, glacial or permafrost							
				Noachian or pre-Noachian bedrock units	●	●	?				11, 0
				Outcrops with remnant magnetization	○	○					0, 11
				Primary, secondary, and basin-forming impact deposits						●	1, 0
				Structural features with regional or global context	●	●		●		●	13, 0
				Diversity of aeolian sediments and/or landforms						●	1, 0

Key	
●	Yes
○	Partial Support or Debated
	No
?	Indeterminate



# Resource ROI(s) Rubric

1<sup>st</sup> EZ Workshop for Human Missions to Mars

Site Factors				RW01-04	RISUR1	LZ01	ISRU2						EZ SUM		
ISRU and Civil Engineering Criteria	Engineering	Meets First Order Criteria (Latitude, Elevation, Thermal Inertia)												1	
	Water Resource	Threshold	AND/ OR	Potential for ice or ice/regolith mix			●								
				Potential for hydrated minerals	●									4	
			Quantity for substantial production	●									4		
			Potential to be minable by highly automated systems	●									4		
			Located less than 3 km from processing equipment site	●									1		
			Located no more than 3 meters below the surface	●									4		
			Accessible by automated systems	●									4		
			Qualifying	Potential for multiple sources of ice, ice/regolith mix <b>and</b> hydrated minerals											
	Distance to resource location can be >5 km	●											3		
	Route to resource location must be (plausibly) traversable	●											3		
	Civil Engineering	Threshold	~50 sq km region of flat and stable terrain with sparse rock distribution												1
			1–10 km length scale: <10°												1
			Located within 5 km of landing site location												1
		Qualifying	Located in the northern hemisphere												0, 1
			Evidence of abundant cobble sized or smaller rocks and bulk, loose regolith												1
	Utilitarian terrain features												1		
	Food Production	Qualifying	Low latitude												1
			No local terrain feature(s) that could shadow light collection facilities												1
			Access to water												1
			Access to dark, minimally altered basaltic sands												1
	Metal/Silicon Resource	Threshold	Potential for metal/silicon												2
			Potential to be minable by highly automated systems												2
			Located less than 3 km from processing equipment site												2
			Located no more than 3 meters below the surface												2
			Accessible by automated systems												
		Qualifying	Potential for multiple sources of metals/silicon												2
			Distance to resource location can be >5 km												2
			Route to resource location must be (plausibly) traversable												2

Key	
●	Yes
○	Partial Support or Debated
	No
?	Indeterminate

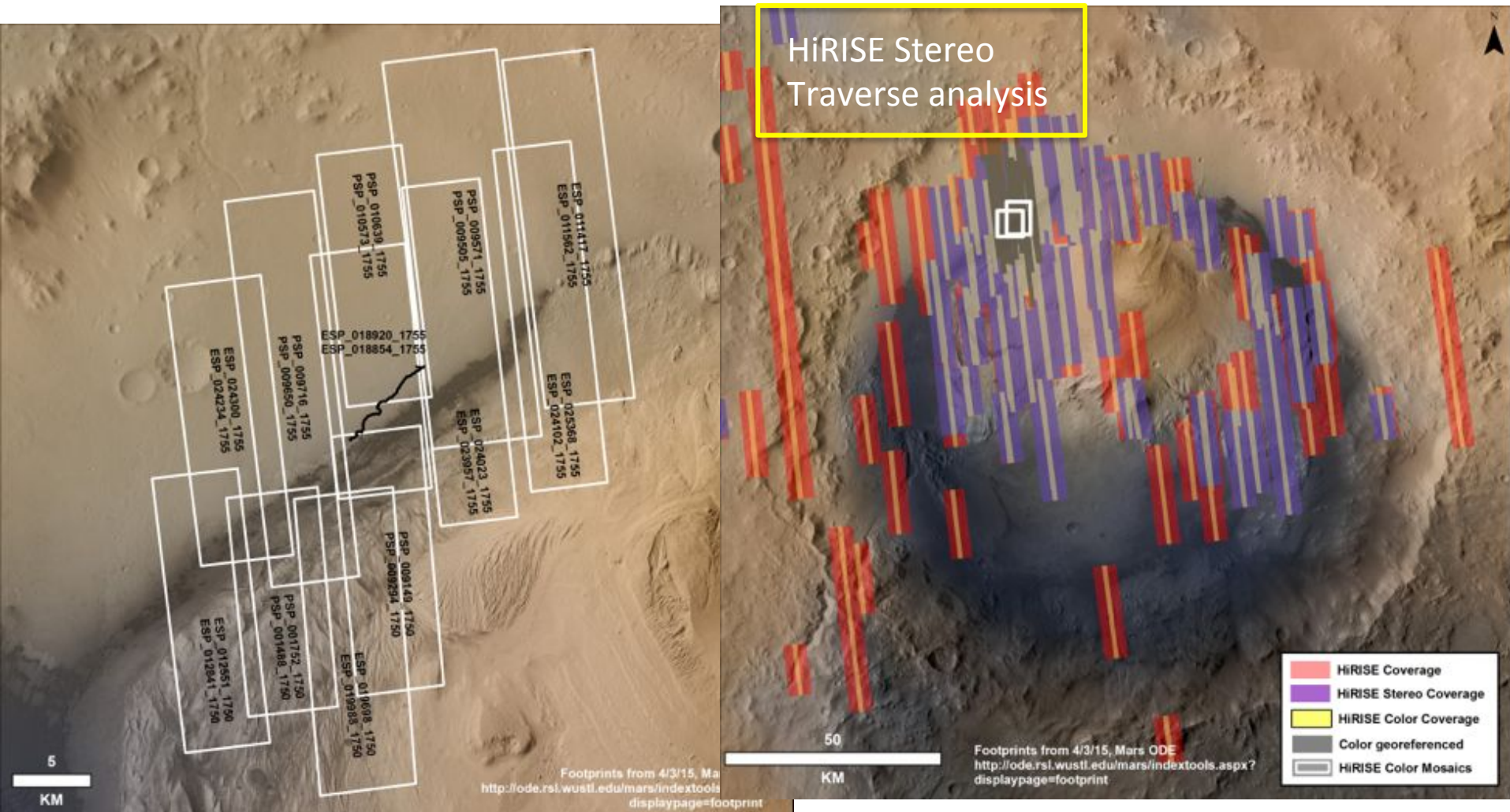


# DATA NEEDS



# HiRISE Coverage Needed for Long Distance Traverse Analysis

1<sup>st</sup> EZ Workshop for Human Missions to Mars





# Highest Priority EZ Data Needs

1<sup>st</sup> EZ Workshop for Human Missions to Mars

- HiRISE Stereo of traverse exiting Gale crater to the north and northwest via Peace Vallis.
- CTX and/or HiRISE Stereo along traverse route outside Gale crater to potential volcanic unit.

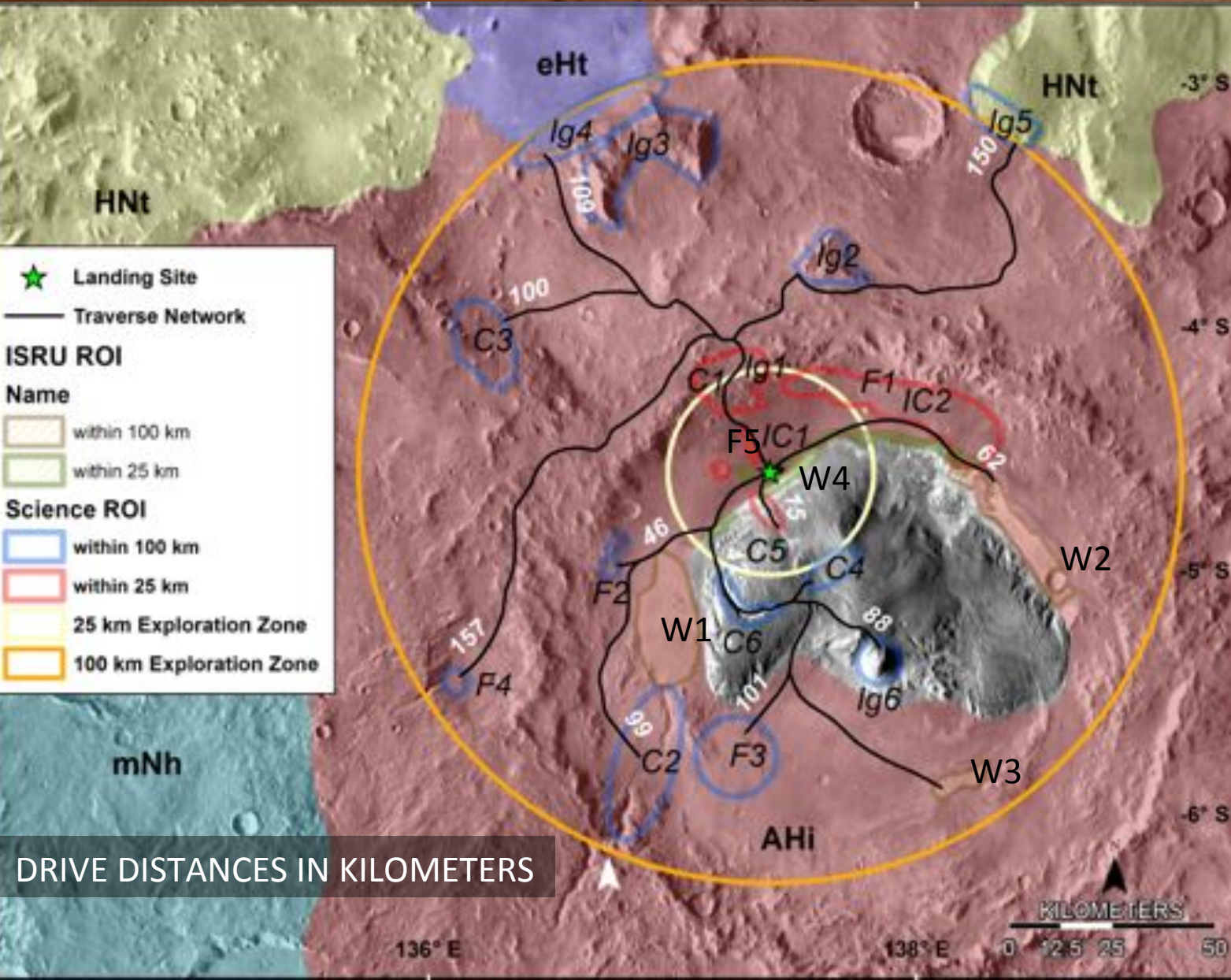


[in order of priority: addressing threshold first, then qualifying]

## **SCIENCE ROIs**



# Gale Crater Exploration Zone



For Human Missions to Mars

## LZ Coordinates

Longitude (E)

137.42009295°

Latitude (S)

4.59310427°

Elevation (MOLA)

-4497.77

Easting

8145534.27 m

Northing

-272254.70 m

## Science ROI

C = Channel

F = Fan

Ig = Igneous

IC = Inverted

Channels

## ISRU ROI

W = Water ROI



# Science ROI: Known Organics

## SROrg01

1<sup>st</sup> EZ Workshop for Human Missions to Mars



Cumberland drill hole site in Yellowknife Bay in Gale crater. The ONLY confirmed presence of insitu martian organics. Image NASA/JPL-Caltech/MSSS/

Gale Crater EZ

ROI, Latitude, Longitude

- SROrg01

137.44909593, -4.58951574, -4520.029 m

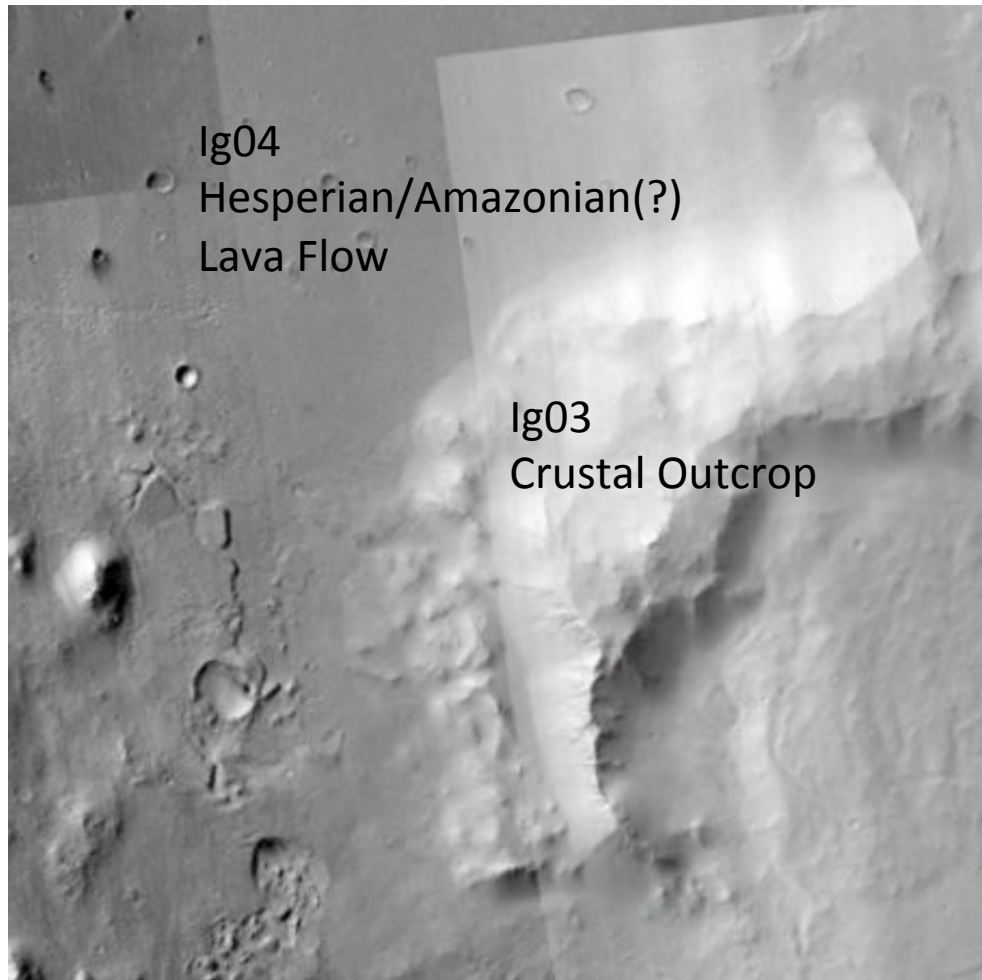
- **Martian Organics: CONFIRMED**
- **Ancient Habitable Environment: CONFIRMED**
- **Aqueous Processes: Lake environment**
- **Goal: access organics below ~ 4 m for best preservation**



# Science ROI: Dateable Rocks

## SRIg01-06

1<sup>st</sup> EZ Workshop for Human Missions to Mars



Dissected Noachian terrain (Ig03)  
and Amazonian (?) lava flow (Ig04)

Gale Crater EZ

### ROI, Latitude, Longitude

- SRIg01  
-4.253, 137.2109, -3514 m
  - SRIg02  
-4.62615916, 137.40954931, -4489.232 m
  - SRIg03  
-3.2417, 136.8176, -643 m
  - SRIg04  
-3.3081, 136.4805, -2634 m
  - SRIg05  
-3.1855, 138.2886, -2253 m
  - SRIg06  
-5.3715, 137.8544, 731 m
- Dateable Crustal Rocks
  - Goal: obtain inplace igneous rocks from Noachian crust and Amazonian lava flow.



# Science ROI: Dateable Igneous Rocks

## SRIg07-11

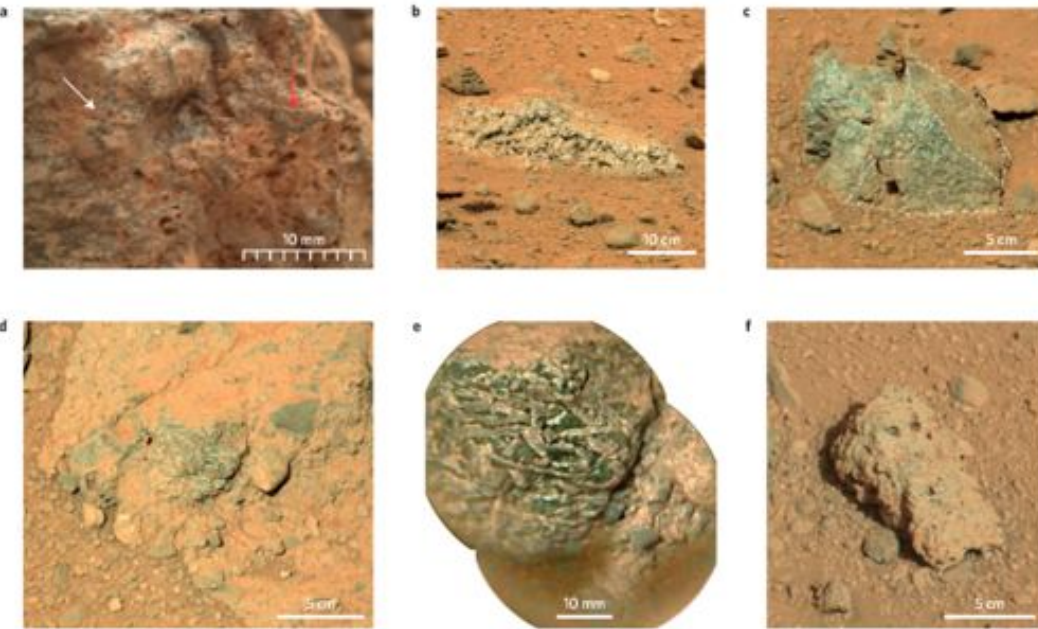
1<sup>st</sup> EZ Workshop for Human Missions to Mars



### ROI, Latitude, Longitude

- SRIg07 Becraft (lat,lon)  
-4.6152702, 137.42078452, -4498.119 m
- SRIg08 Clinton (lat,lon)  
-4.62615916, 137.40954931, -4489.232 m
- SRIg09 Harrison (lat,lon)  
-4.62617745, 137.40955945, -4489.114 m
- SRIg10 Little Wind River (lat,lon)  
-4.59700021, 137.4369205, -4500.566 m
- SRIg11 Sparkle (lat,lon)  
-4.62620416, 137.40957441, -4489.030 m

- Dateable Igneous (Crustal ?) Rocks
- All located close to LZ
- Goal: obtain Noachian aged samples likely from the crater rim.

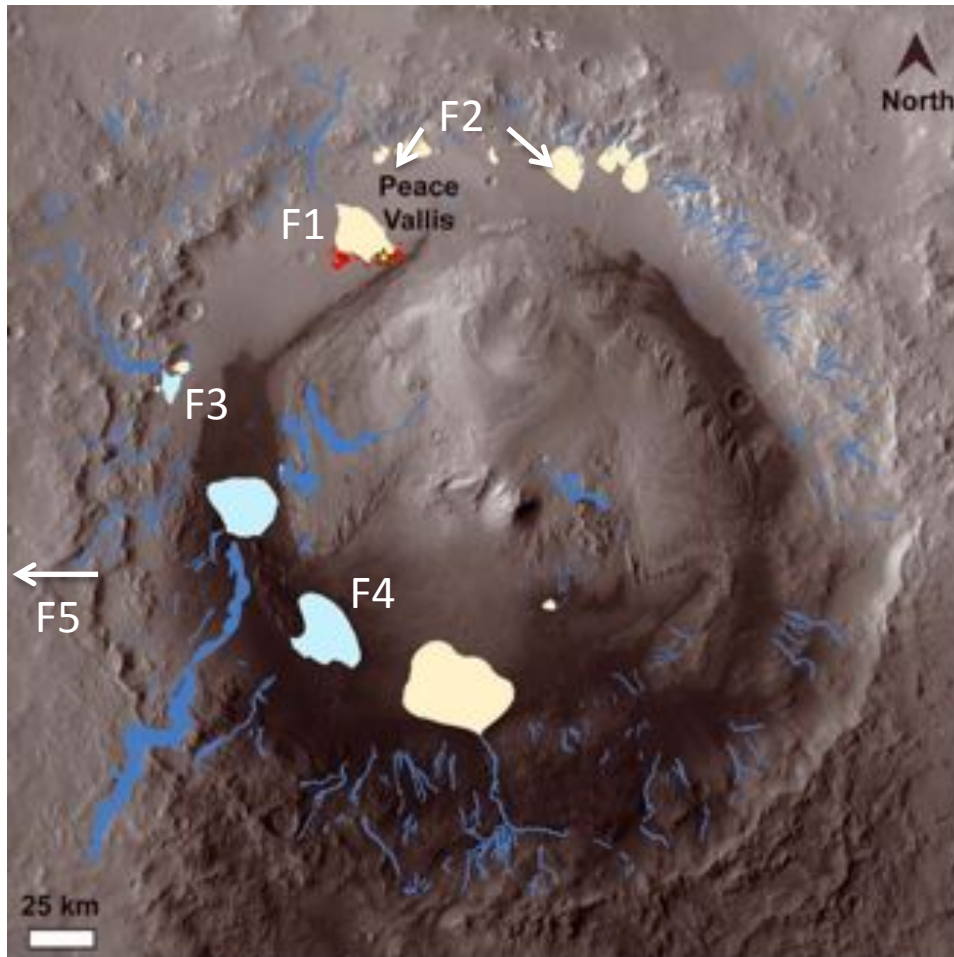


Examples of various igneous rocks found along Curiosity's traverse, from Sautter et al., 2015. These rocks are expected to be from a local source within the crater, possibly from the rim.



# Science ROI: Alluvial Fans and Potential Deltas SRF01-05

1<sup>st</sup> EZ Workshop for Human Missions to Mars



Alluvial fans and potential deltas  
from Palucis et al., 2013

## ROI, Latitude, Longitude

- SRF01 Peace Vallis Alluvial Fan  
-4.4549, 137.3476, -4464 m
  - SRF02 Alluvial fans NE of Peace Vallis  
-4.2852, 137.7909, -3995m
  - SRF03 small delta (?) on western rim  
-5.0038, 136.7512, -3917 m
  - SRF04 'pancake delta' N of Farrah Vallis  
-5.7801, 137.2926, -3268 m
  - SRF05 fan/delta just W of crater rim  
-5.4686, 136.1128, -1474 m
- 
- Aqueous Processes
  - Potential Habitable Environments
  - Potential Organic Preservation
  - Investigate the diverse water level history in Gale/region



# Science ROI: Inverted Channels

## SRIC01-02

1<sup>st</sup> EZ Workshop for Human Missions to Mars



Peace Vallis inverted Channel  
NASA/JPL-Caltech/UofA

ROI, Latitude, Longitude

- SRIC01  
-4.4549, 137.3476, -4464 m
- SRIC02  
-4.2852, 137.7909, -3995m

- Aqueous Processes



# Science ROI: Channels

## SRC01-06

1<sup>st</sup> EZ Workshop for Human Missions to Mars

### ROI, Latitude, Longitude

- SRC01 Peace Vallis  
-4.253, 137.2109, -3514 m
  - SRC02 Farrah Vallis  
136.8125, -5.9589, -2302 m
  - SRC03 unnamed channel NW of rim  
-4.0385, 136.2456, -2159 m
  - SRC04 exhumed channel under upper mound  
-5.1011, 137.5155, -2012 m
  - SRC05 Lower Mt. Sharp  
-4.7876, 137.4095, -3650 m
  - SRC06 western mound  
-5.1417, 137.2671, -2998 m
- 
- Stratigraphic context (late Noachian to Hesperian/Amazonian)
  - Aqueous Processes  
(in cross-section)
  - Dateable Surfaces (?)



Channel cutting western part of lower mound.

NASA/JPL-Caltech/UofA

Gale Crater EZ



[in order of priority: addressing threshold first, then qualifying]

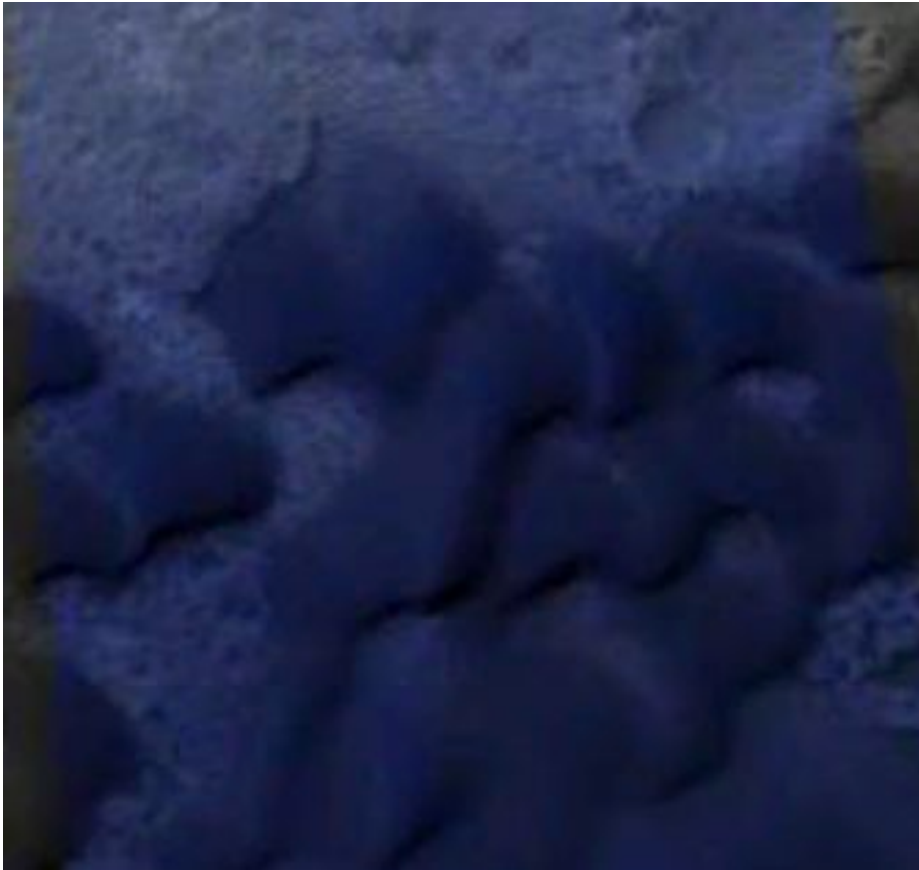
## **RESOURCE ROIs**



# H<sub>2</sub>O ISRU Resource

## ISRUW01-04

1<sup>st</sup> EZ Workshop for Human Missions to Mars



Barchan dunes of Bagnold dunefield.  
NASA/JPL-Caltech/UofA

### ROI, Latitude, Longitude

- ISRUW01  
-4.253, 137.2109, -3514 m
  - ISRUW02  
136.8125, -5.9589, -2302 m
  - ISRUW03  
136.8125, -5.9589, -2302 m
  - ISRUW04  
136.8125, -5.9589, -2302 m
- 
- Adsorbed H<sub>2</sub>O at 2-3 wt%. Estimated 10<sup>4</sup>-10<sup>6</sup> MT in dunes, ripples, loose regolith
  - H<sub>2</sub>O in amorphous component, 3-6 wt%
  - Hydrated Phyllosilicates, 2-3 wt%
  - Excellent building material for sand bags/radiation protection



# Metal and Silica ISRU Resources

## ISRU01-02

1<sup>st</sup> EZ Workshop for Human Missions to Mars



ROI, Latitude, Longitude

- ISRU01 (Al, Si)  
-4.253, 137.2109, -3514 m
- ISRU02 (Fe)  
Yellowknife Bay  
-4.58951574, 137.44909593,  
-4520.029 m

- ISRU01 rocks on average have  
~12 wt% Al and ~55-60 wt% Si.  
ISRU02 rocks, ~18 wt% Fe.
- ISRU1
  - **300 kg/m<sup>3</sup> Al**
  - **1375-1500 kg/m<sup>3</sup> Si**
- ISRU2
  - **450 kg/m<sup>3</sup> Fe**

Outcrop 'Link', a matrix-supported conglomerate with relatively high Al and Fe content. Exposures near surface form a natural 'pavement' and ~1 cm sized clasts.

NASA/JPL-Caltech/MSSS

Gale Crater EZ